

<sup>1</sup>Genetic relationship among komatiites and associated basalts  
<sup>2</sup>in the Badampahar greenstone Belt (3.25–3.10 Ga),  
<sup>3</sup>Singhbhum Craton, Eastern India

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## 11Abstract

12The ultramafic-mafic volcanic rocks of Archean greenstone belts are important archives for  
13lithospheric and asthenospheric processes of the early Earth. Despite decades of research on  
14this context, many issues still remain unsolved. For example, the process of komatiite magma  
15genesis and the genetic relationship among komatiites, komatiitic basalts and tholeiitic basalts  
16in Archean greenstone belts are not clearly understood. The metavolcanic rocks of the  
17Badampahar greenstone belt (BGB), Singhbhum Craton are studied by major-trace element  
18geochemistry to address the said problems and better understand the evolution of melts in  
19Archean lithosphere. Our research suggests that the protoliths of the metavolcanic rocks were  
20komatiites (both Al –depleted and –undepleted), komatiitic basalts and tholeiitic basalts. The  
21Al–heavy rare earth element (HREE) depleted komatiites were formed by moderate degree  
22mantle melting at a higher depth and the Al–HREE undepleted komatiites are products of  
23moderate to high degree mantle melting at a shallower depth. The melting–assimilation–  
24fractional crystallization modelling result shows that komatiitic basalts were generated from  
25Al–undepleted komatiites, and tholeiitic basalts were generated from evolved komatiitic  
26basalts by assimilation and fractional crystallization processes. The older age limit of the  
27BGB is determined to be 3.25 Ga. and the basement of sedimentation and volcanism was  
28composed of plutonic felsic rocks.

29**Keywords:** komatiite; komatiitic basalt; tholeiitic basalt; greenstone belt; crustal  
30assimilation.

### 311. Introduction

32 Komatiites together with associated basalts can provide insight into the  
33 physicochemical conditions of mantle and magma generation history of the early Earth  
34 (Hofmann, 1997; Arndt et al., 2008; Condie, 2015, Condie et al., 2016). Extensive studies  
35 have been done on these Archean ultramafic–mafic rocks for the last four decades (e.g.,  
36 Barnes et al., 1988; Jayananda et al., 2008; Stiegler et al., 2010; Dostal and Mueller, 2013;  
37 Konnunaho et al., 2013; Mole et al., 2014; Verma et al., 2017 and references therein).  
38 However, many petrogenetic aspects of this rock assemblage still lack transparency. For  
39 example, komatiites and komatiitic basalts though have textural–geochemical (major  
40 elements) similarities and occur together in greenstone belts, the genetic relationship between  
41 them is not clear. It is a matter of debate whether komatiitic basalts were generated by  
42 fractional crystallization of komatiites or as primary mantle melt (Arndt et al., 2008 and  
43 references therein). A number of studies (e.g. Francis and Hynes, 1979; Arndt and Nesbitt,  
44 1982; and Cattell, 1987) suggest that fractional crystallization and assimilation of komatiites  
45 with granitoid rocks might have led to the generation of komatiitic basalts. However, most  
46 other studies (e.g. Nesbitt and Sun, 1976; Arndt and Nesbitt, 1984; Arndt et al., 1997; Hanski  
47 et al., 2001; Dostal and Mueller, 2013; Verma et al., 2017) consider diverse origins of  
48 komatiites and associated basalts. The relationship of tholeiitic basalts with komatiites and  
49 komatiitic basalts is also uncertain as tholeiitic basalts have very different and evolved  
50 chemical compositions (Arndt, 1991; Condie, 1994). The interlayered association of different  
51 volcanic rocks in Archean greenstone belts indicating their formation simultaneously arises  
52 questions about their genetic connection. The present paper addresses the aforesaid  
53 geological problems with the geochemical evidences gathered from the ‘meta’ ultramafic-  
54 mafic volcanic rocks of Badampahar greenstone belt (BGB), Singhbhum Craton, Eastern  
55 India.

56 A number of studies (e.g. Sahu and Mukerjee, 2001; Chaudhuri et al., 2015, 2017)  
57 have previously reported presence of 'meta' -komatiites and -basalts from the eastern  
58 greenstone belts of Singhbhum Craton and the studies were mostly concentrated upon  
59 petrogenesis of komatiites. However, none of the studies had attempted to investigate the  
60 genetic connection among komatiites and associated basalts of the studied area. Absolute  
61 geochronology of the studied greenstone rocks by direct radiometric age dating is absent.  
62 Saha (1994) suggested an indirect age ca. 3.3-3.1 Ga of the rocks defined by their  
63 stratigraphic relationship with associated granitoid rocks. Hence, a geochronological re-  
64 evaluation of the BGB rocks is required alongside the geochemical study of metavolcanic  
65 rocks.

66 The metavolcanic rocks of BGB are studied by major and trace element geochemistry with  
67 the following objectives, (i) better understanding the genesis of komatiites, and melt  
68 generation processes in Archean mantle, (ii) exploring the genetic connection among  
69 komatiites and associated basalts, and (iii) modelling the geochemical evolution of melts in  
70 Archean lithospheric reservoirs. We have also studied detrital zircon geochronology of a  
71 metaconglomerate (occurring at lower part of stratigraphy) with the objective to know the age  
72 and basement rocks of the studied greenstone belt which is relevant to this study.

## 732. Geological setting

74 The Singhbhum Craton (Fig.1A) of the Eastern India is bounded by the Singhbhum  
75 Mobile Belt to the north, the Bastar Craton to the west, Eastern Ghats mobile belt to the south  
76 and southeast (Saha, 1994). It principally comprises Paleo- and Meso- Archean supracrustal  
77 rocks belonging to Older Metamorphic Group (OMG), Older Metamorphic Tonalitic Gneiss  
78 (OMTG), Singhbhum Granite and equivalents, and Iron Ore Group (Mukhopadhyay, 2001;  
79 Table 1). The central part of Singhbhum Craton is principally occupied by granitoid rocks of

80 different generation belonging to Singhbhum Granite, whereas the rocks of other group occur  
81 mostly at the peripheral part (Saha, 1994; Mishra et al., 1999). The OMG is the oldest rock  
82 group in the craton and it encompasses amphibolite facies supracrustal assemblage of pelitic  
83 schists, quartz–magnetite–cummingtonite schists, calc–gneiss and amphibolites (Saha, 1994;  
84 Mukhopadhyay, 2001). The depositional time span of OMG rocks is considered between 3.50  
85 Ga. and 3.45 Ga, (U-Pb date of zircon; Goswami et al., 1995; Mishra et al. 1999). A number  
86 of studies (e.g. Saha, 1994; Gowshami et al., 1995; Upadhyay et al., 2014) reported that the  
87 OMG rocks occur as enclaves within younger tonalities-trondhjemites-granodiorites (TTGs)  
88 belonging to OMTG. The first phase of the TTGs was emplaced around 3.45 Ga (U-Pb date  
89 of zircon; Mishra et al., 1999; Acharyya et al., 2010; and Upadhyay et al., 2014). Upadhyay  
90 et al. (2014) reported another younger phase of it emplaced around 3.32 Ga (U-Pb date of  
91 zircon). The OMG and OMTG occur as enclave within Singhbhum Granite, and the biggest  
92 exposure of them is located at the western part of the craton (Fig. 1B), near village Champua  
93 (Mukhopadhyay, 2001). The Singhbhum Granite mainly comprises granodiorites,  
94 trondhjemites and granites (Saha, 1994). Upadhyay et al. (2014) reported that three major  
95 phases of Singhbhum Granite Type–A plutonic intrusions occurred synchronously with  
96 OMTG emplacement during 3.44 Ga, 3.35 Ga and 3.32 Ga (U-Pb date of zircon). The  
97 plutonic intrusions of Singhbhum Granite Type–B (ca. 3.12 Ga; Pb-Pb date by Ghosh et al.  
98 1996) and equivalent rocks such as Bonai granite (ca. 3.16 Ga; U-Pb date of zircon by  
99 Sengupta et al., 1996) and Mayurbhanj granite (ca. 3.09 Ga; U-Pb date of zircon by Mishra et  
100 al., 1999) happened after deposition and metamorphism of the Iron Ore Group rocks (Saha,  
101 1994).

102       At the peripheral part, the Singhbhum Craton is flanked by several greenstone belt  
103 successions (belonging to the Iron Ore Group) which are exposed near the Jamda–Koirā  
104 region in the west, the Gorumahisani–Badampahar region in the east and the Tomka–Daitari

105 region in the south. Saha (1994) interpreted them to be formed in between 3.3 Ga and 3.1 Ga  
106 by their stratigraphic relationship with associated granitoid rocks. The volcano-sedimentary  
107 greenstone sequences are mostly consisting of banded iron formations, mafic-felsic volcanic  
108 rocks, shale, quartzite, chert and carbonates (Saha 1994; Mukhopadhyay, 2001). The  
109 Badampahar greenstone belt (BGB) occurs at the southern part of the Gorumahishani–  
110 Badampahar region (eastern Iron Ore Group; Fig. 1B). The rocks of the BGB are  
111 metamorphosed from greenschist to lower amphibolite facies (Saha, 1994; Ghosh and  
112 Baidya, 2017a).

### 1133. Stratigraphy and sampling

114 The BGB majorly comprises interbedded sequence of metavolcanic rocks, quartzite,  
115 banded iron formations, chert and phyllite along with minor metavolcanic clastics and schists  
116 (Fig. 1C). Detailed structural study reveals at least four major phases of deformation (Ghosh  
117 and Baidya, 2017a). The first deformation phase ( $D_1$ ) developed the regional Badampahar  
118 isoclinal fold ( $F_1$ ) plunging  $63^\circ$  towards NNE and foliation  $S_1$ . The second deformation ( $D_2$ )  
119 caused extensive shearing and folding ( $F_2$ ) with the development of  $S_2$  foliation.  $S_2$  dips from  
120  $75^\circ$  to  $85^\circ$  towards NW. Isoclinal to tight  $F_2$  folds plunge from  $25^\circ$  to  $63^\circ$  towards NE or SW  
121 (Fig. 1C: near Jashipur). The third deformation phase ( $D_3$ ) caused open type folds ( $F_3$ )  
122 plunging from  $40^\circ$  to  $75^\circ$  towards NW. Faults and joints were developed subsequently. Three  
123 sets of fault trending NW-SE, E-W and N-S are interpreted from field and satellite imagery.  
124 Three sets of penetrative joints are present, dipping  $80^\circ - 85^\circ$  towards ESE,  $70^\circ - 80^\circ$  towards  
125 SE to SSE and vertical joints trending N-S.

126 Stratigraphic sections were prepared based on mapping along AB, CD, EF, GH, IJ,  
127 KM, LM, OP and NP on the limbs of regional isoclinal fold (Fig. 1C). Cross stratification and  
128 normal graded bedding in quartzites and pillows in metavolcanic rocks confirmed the local

129younging direction of the sequences, and, together with the broad  $F_1$  and  $F_2$  fold structure,  
130allowed regional correlations. The stratigraphic sections are thoroughly described in  
131Supplementary material 1. In summary, most stratigraphic sections of the BGB show that the  
132lower and upper –part of the sequences (Fig. A1, A3 and A5, Supplementary material 1) are  
133occupied by metavolcanic rocks (Fig. A2a-A2e, supplementary material 1) along with a few  
134sedimentary interbeddings. Section EF and GH exceptionally do not comprise a lower and an  
135upper metavolcanic rock horizons. The middle part of all stratigraphic sections are mostly  
136occupied by sedimentary rocks. The metavolcanic rock horizons of the middle part are  
137relatively thin (10–20 m) and interlayered with sedimentary rocks (Fig. A2f, supplementary  
138material 1). Correlation of the stratigraphic sections in the BGB led to categorise the  
139metavolcanic rocks in three groups, namely basal–, middle– and upper– metavolcanic rocks.

140       A total twenty three fresh metavolcanic rock samples devoid of any vesicle/amygdule,  
141visible weathering and hydrothermal vein are selected for bulk geochemical analyses. Among  
142them, eleven samples of the basal–metavolcanic rocks from section AB, NP and KM, six  
143samples of the middle–metavolcanic rocks from section EF and GH, six samples of the  
144upper–metavolcanic rocks from section AB, LM and IJ are selected for analysis  
145(Supplementary material 1). We collected the metaconglomerate sample (A33) from section  
146IJ for petrography and detrital zircon dating.

#### 1474. Analytical techniques

148Electron probe micro analyser (EPMA) was used to determine the chemical compositions of  
149minerals present in the studied metavolcanic rocks. EPM analysis was done using a  
150‘CAMECA SX Five’ model in the laboratory of Indian School of Mines, Dhanbad, India with  
151an accelerating voltage 15kV, current 12nA and 1  $\mu\text{m}$  beam diameter. Standards were run  
152before and after analysis to determine the analytical error. The analytical inaccuracies and

153uncertainties of the analysis are negligible and as follows: <0.1% for Cr; <1% for P, Al, Si,  
154K, Ca, Fe, Mg, Ti, V; <2% for Mn and Na.

155 The metavolcanic rock samples were pulverized for major, trace and rare earth  
156elements (REEs) analyses at the laboratory of Australian Laboratory Services Pty Ltd., Perth,  
157Australia. The samples were digested and analysed for major oxides in inductively coupled  
158plasma–atomic emission spectroscopy (ICP–AES) following the methods of Murray et al.  
159(2000). For trace and REEs, the samples were acid digested and analysed using inductively  
160coupled plasma–mass spectroscopy (ICP–MS) following the methods of Qing et al. (2003).  
161The laboratory standard reference materials namely OREAS146, GGC–8, GGC–10,  
162GBM908–10 and MRGEO08 are analysed along with each batch for accuracy and precision.  
163The analytical uncertainties for major and trace elements are <4% and <10%. The complete  
164dataset with sample location is available in Table A3 (Supplementary material 5).

165 Zircon grains were separated from the metaconglomerate by conventional magnetic  
166and heavy liquid separation after crushing and grinding. The separated zircon grains were  
167mounted in Araldite on glass slides and polished. The U–Th–Pb isotopic analyses were done  
168on an Agilent 7700x ICP–MS coupled to a NWR193 excimer laser system at the London  
169Geochronology Centre. Zircons were dated ‘blindly’ so as to avoid selection bias (Garzanti et  
170al., 2018), and were only subjected to cathodo–luminescence imaging after U–Pb analysis.  
171The laser was operated at 20Hz pulse repetition rate, >2 J/cm<sup>2</sup> energy density, 25 µm laser  
172pit diameter and 20s duration. Plešovice zircon and NIST SRM612 glass were analysed along  
173with the studied zircons for age and concentration standardization. Linear fit through the  
174standards yielded a fit within 1% for the <sup>206</sup>Pb/<sup>238</sup>U ratios. Data reduction was performed using  
175Glitter 4.4 (Griffin et al., 2008). The complete dataset with sample location is available in  
176Table A2 (Supplementary material 4). Further post processing, including concordia analysis  
177and kernel density estimation were done with IsoplotR version 1.1 (Vermeesch, 2018).



## 1785. Detrital zircon geochronology of metaconglomerate

179        The metaconglomerate horizon in stratigraphic section IJ (250m) overlies a quartzite  
180horizon and grades upward to another quartzite (Fig. A5; Supplementary material 1). The  
181metaconglomerate is clast supported with minor quartz rich matrix (< 10 modal %).  
182Secondary chloride–sericite–goethite are present along the clast boundaries and inter–  
183granular spaces (Fig. 2A, 2B). The clasts are well rounded and ellipsoidal with 1–2 cm long  
184axes. The long axes of the clasts are oriented along the regional first generation foliation  
185plane ( $S_1$ ) corresponding to  $D_1/F_1$  (Fig. 2A). The clasts consist mostly of quartz aggregates  
186with minor K–feldspar, zircon and apatite. The quartz grains exhibit a crude interlocking  
187texture with mostly straight and sharp grain boundaries (Fig. 2B, 2C and 2D). However, some  
188grain boundaries are weakly diffused by pressure solution. The zircon grains of various size  
189(20–200 $\mu$ m long axis) and shapes are present as inclusion within quartz.

190        The CL images of the analysed zircon grains exhibit distinct oscillatory growth  
191zoning by alternating high and low CL response to the zones (Fig. 2E–2L). Plotting the U–Pb  
192data for sample A33 on a Wetherill concordia diagram (Fig. 3A, 3B) defines a discordia trend  
193with a  $\sim$ 875 Ma lower intercept which indicates partial Pb–loss during an unidentified  
194Neoproterozoic tectonic event. The Kernel Density Estimate (KDE, Vermeesch, 2012) of the  
195concordant ages reveal two distinct components at  $\sim$ 3.30 and  $\sim$ 3.45 Ga., respectively (Fig.  
1963C). The age of the youngest zircon grain of the assemblage is 3.25 Ga.

## 1976. Petrography and geochemistry of the metavolcanic rocks

### 1986.1. Petrography

199        The rocks of BGB have endured multiple events of deformation and metamorphism.  
200Hence, original igneous texture is almost obscured and initial minerals are replaced by

201secondary metamorphic minerals. However, the rocks could be classified into three groups  
202based on metamorphic mineralogy and texture. The first group of rocks primarily consists  
203high-Mg silicate minerals such as serpentine and talc (65-80 modal %) with the ubiquitous  
204presence of tremolite (10-15 modal %) and magnetite (15-20%). Secondary minerals such as  
205serpentine, talc and tremolite pseudomorph after original igneous minerals (e.g. olivine,  
206orthopyroxene and clinopyroxenes) in these rocks (Fig. 4A and 4B). The coarse (200-  
2071000 $\mu$ m) pseudomorphic- and magnetite- crystals occur within the fine (20-100 $\mu$ m)  
208serpentine-talc rich groundmass (Fig. 4A and 4B). Olivine and pyroxenes are absent in these  
209rocks due to their complete alteration to secondary minerals. Relict of spinifex texture can be  
210identified in some samples by radiating long (3-5 mm) skeletal/acicular crystals of magnetite  
211and talc-serpentine pseudomorphs after olivine/orthopyroxene (Fig. 4C). The second group of  
212rocks consists tremolite-actinolite (50-70 modal %), magnetite (10-15 modal %) and  
213serpentine-talc (20-30 modal %) as major minerals and plagioclase (<2-10 modal %) as  
214minor mineral (Fig.4C and 4D). Whereas, the third group of rocks are primarily consist of  
215hornblende (65-85 modal %) and plagioclase (10-25 modal %) and magnetite (5-10 modal %)  
216(Fig. 4E and 4F). Ilmenite, titanite, zircon, apatite and sulphides are present as accessory  
217minerals in these rocks. The amphibole porphyroblasts varies in size from very fine (100 $\mu$ m)  
218to coarse (2000 $\mu$ m) crystals. Most samples show strong metamorphic fabric defined by  
219amphibole grain alignment along regional first generation foliation plane (Fig. 4E and A2d,  
220supplementary material 1). Post-kinematic amphibole porphyroblasts cut across this fabric.  
221Original igneous minerals and textures are though absent, some of the rocks contain initial  
222ophitic and subophitic textural outlines preserved in aggregates of plagioclase and amphibole  
223crystals (Fig. 4D and 4F). The amphiboles are possibly pseudomorphs after clinopyroxene.  
224The plagioclase crystal varies in compositions (albite-andesine) and in size (from 100 $\mu$ m to  
2251000 $\mu$ m) both within sample and between samples.

## 2266.2. Effect of metamorphic and post–metamorphic events

227        The rocks of BGB have experienced multiple phases of deformation, up to lower  
228 amphibolite facies of metamorphism and some hydrothermal events (Ghosh and Baidya,  
229 2017a) which may mask the original geochemical characteristics. Therefore, it is important to  
230 detect those altered rocks by petrography and geochemical filters before any interpretation.  
231 The analysed rocks are examined under microscope, and most of the samples are free from  
232 any visible alteration, amygdule/vesicle and hydrothermal veins. The loss in ignition (LOI)  
233 values indicate presence of volatiles in rock (as carbonate minerals) and may indicate degree  
234 of alteration. Most of the studied rocks have low LOI content ranging from 0.79 to 2.25 wt.  
235 %, but some of them (sample E35, D84, D79, E32, E24, D118 and D120) have little higher  
236 LOI values ranging from 4.79 to 10.68 wt. %. However, the carbon concentration (<0.1 wt.  
237 %) is negligible in all of the samples (Table A3, supplementary material 5). This rules out the  
238 possibility of element incorporation from external source by carbonate alteration. The high  
239 LOI values are due to the hydrous silicate mineralogy (e.g., hornblende, talc and serpentine)  
240 of the studied rocks. Moreover, the LOI values show little or no correlation with Nb/La and  
241 Th/La ratios (as La is considered less mobile in alteration relative to Nb and Th) indicating  
242 that the elemental ratios have not been affected by the increase of volatile content (Fig. A6,  
243 Supplementary material 3). Zirconium is one of the least mobile elements and it may be used  
244 as a reference to test the mobility of other elements. The Zr vs. major oxides ( $\text{Fe}_2\text{O}_3$ , MgO  
245 and  $\text{TiO}_2$ ), high field strength elements (HFSEs; Th and Nb) and REEs (La, Sm and Yb)  
246 diagrams (Fig. A7 and A8, Supplementary material 3) show mostly coherent trend or limited  
247 scatter (exception for sample D79), implying immobility of the elements during metamorphic  
248 and post–metamorphic events. The NMORB (Sun and McDonough, 1989) normalized multi–  
249 element spider diagram (Fig. A9, supplementary material 3) shows sub–parallel/parallel REE  
250 and HFSE (Hf, Zr, Ti, Ta and Nb) patterns of the metavolcanic rocks (exception for D79)

251 indicating minor or no mobility during metamorphism. This fact is also supported by near  
252 zero  $Ce_{Cl}$  and  $Eu_{Cl}$  (CI chondrite normalized) anomalies of the most analysed samples (Table  
253) as considerable amounts of Ce and Eu anomalies indicate alteration by weathering and  
254 hydrothermal processes. The LILE are considered highly mobile elements during  
255 metamorphism and alteration. Most of the metavolcanic rock samples show irregular and  
256 enriched LILE patterns which are expected in deformed–metamorphosed Archean greenstone  
257 rocks. The metavolcanic rock samples excluding sample D79 are considered for  
258 geochemistry and mainly the less mobile elements such as transition metals, HFSEs and  
259 REEs are selected for interpretation.

### 260 6.3. Major and trace element compositions

261 The metavolcanic rocks are characterized by variable but significant contents of  $SiO_2$   
262 (38.4–57.4 wt. %),  $Al_2O_3$  (1.41–13.15 wt. %),  $CaO$  (1.52–11.9 wt. %),  $Fe_2O_3^T$  (8.47–17.3 wt.  
263 %),  $TiO_2$  (0.07–1.91 wt. %) and  $MgO$  (3.05–32.7 wt. %) and  $Mg\#$  varies between 15 and 75  
264 (Table A3, Supplementary material 5). They dominantly occupy the area of basalt to basaltic  
265 andesite protolith in  $SiO_2$  vs.  $K_2O+Na_2O$  plot (TAS; after Le Bas et al., 1986; Fig. 5A) and  
266  $Nb/Y$  vs.  $Zr/Ti$  plot (after Pearce, 1996; Fig. 5B). The subalkaline nature of these rocks are  
267 also evident in the TAS diagram due to their low alkali element concentrations (Irvine and  
268 Baragar, 1971).

269 The metavolcanic rocks based on  $MgO$  contents (and petrographic study) can be  
270 divided into three groups those will be later determined to be komatiites, komatiitic basalts  
271 and tholeiitic basalts. Group–I rocks (Sample E35, D84, E32, E25, E24, D118 and D120)  
272 have  $MgO$  values between 19.05 and 32.7 wt. %; group–II rocks (Sample E34, E39, E36,  
273 E37, E40, E44, D11, E38, A10, E23, D123 and A26) have  $MgO$  values between 8.47 and  
274 14.55 wt. %; and group–III rocks (Sample D33, E11, E13, E17, E22 and D111) have  $MgO$

275 values between 3.05 and 6.46 wt. %. These groups also have contrasting SiO<sub>2</sub>, Al<sub>2</sub>O<sub>3</sub>, Ga, Ni,  
 276 Cr, Co, Y and Zr concentrations. For example, the group–I rocks contain lower  
 277 concentrations of SiO<sub>2</sub> (38.4–48.4 wt. %), Al<sub>2</sub>O<sub>3</sub> (1.41–8.01 wt.%), Y (2.7–7.9 ppm), Ga  
 278 (1.5–6.5 ppm), V (24–163 ppm) and  $\Sigma$ REE (3.1–33.98 ppm) and higher concentrations of Cr  
 279 (960–6410 ppm), Ni (777–2340 ppm) and Co (75–125 ppm) than the group–II and group–III  
 280 rocks. Concentrations of Al<sub>2</sub>O<sub>3</sub> (5.73–10.95 wt.%), Y (9.4–17.9 ppm), Ga (8.1–12.2 ppm), V  
 281 (159–279 ppm), Cr (600–1490 ppm), Ni (173–525 ppm), Co (50–77 ppm) and  $\Sigma$ REE (14.28–  
 282 44.27 ppm) in the group–II rocks are intermediate relative to other groups. The group–III  
 283 rocks have higher concentrations of Al<sub>2</sub>O<sub>3</sub> (10.4–13.15 wt.%), Y (16.8–45.9 ppm), Ga (16.4–  
 284 19.1 ppm), V (243–454 ppm) and  $\Sigma$ REE (50.12–134.9 ppm) and lower concentrations of Cr  
 285 (40–380 ppm), Ni (51–114 ppm) and Co (40–51 ppm) than the group–I and group–II rocks  
 286 (Table A3, Supplementary material 5). The group–I and group–II rocks also have lower Zr  
 287 and Nb concentrations (Zr~6–64 ppm and Nb~ 0.2–3.4 ppm) than group–III rocks (Zr~84–  
 288 239 ppm and Nb~4.6–12.4 ppm).

289       The group–I metavolcanic rocks have MgO >18 wt. % which may be present in few  
 290 high Mg volcanic rocks such as, komatiites, picrites, meimechites and boninites. Among  
 291 them, meimechites and komatiites have Na<sub>2</sub>O+K<sub>2</sub>O values (<2 wt. %) less than picrites  
 292 (Maitre et al., 1989). Meimechites have generally high TiO<sub>2</sub> contents (>1 wt. %) and  
 293 boninites have high SiO<sub>2</sub> contents (>53 wt. %) which are absent in komatiites (Le Bas, 2000).  
 294 Boninites also show high LILE concentrations and ‘U’ shaped chondrite normalized REE  
 295 patterns due to MREE depletion (Smithies et al., 2004). The Group–I rocks have Na<sub>2</sub>O+K<sub>2</sub>O  
 296 contents between 0.06 and 0.56 wt. % (Fig. 5A), TiO<sub>2</sub> values between 0.07 and 0.65 wt. %,   
 297 SiO<sub>2</sub> values between 38.4 and 48.4 wt. % and low LILE contents (Table A3, Supplementary  
 298 material 5). These also do not exhibit depleted MREE patterns in chondrite normalization  
 299 (Fig. A9 in supplementary material 3; Table 3). Therefore, protolith of group–I metavolcanic

300rocks was komatiite. These also occupy the komatiite zone in the Mg–Fe+Ti–Al triangular  
301diagram proposed by Jensen (1976), along with group–II rocks at the komatiitic basalt zone  
302and group–III rocks at high Fe tholeiite zone. In MgO vs. TiO<sub>2</sub> diagram (after Arndt et al.,  
3032008), group–I rocks are also plotted at the komatiite zone (Fig. 5D). These have Mg#  
304between 59 and 75, Al<sub>2</sub>O<sub>3</sub>/TiO<sub>2</sub> ratios between 2.43 and 61.6, and CaO/Al<sub>2</sub>O<sub>3</sub> ratios between  
3050.37 and 4.2 (Table 3).

306 Komatiitic basalts are generally distinguished from other basalts by their spatial  
307association with komatiites, textures and chemical compositions. For example, these have  
308lower MgO values (between 18 and 6 wt. %) and higher Al<sub>2</sub>O<sub>3</sub>, Ti and Zr values than  
309komatiites and they plot at the extension of chemical trends defined by komatiites (Arndt et  
310al., 1977; Arndt and Nebitt, 1982; Arndt et al., 2008). The group–II rocks are interbedded  
311with the komatiites (group-I rocks) in the BGB volcano–sedimentary sequences and have  
312lower MgO and higher Al<sub>2</sub>O<sub>3</sub> values than komatiites. These also have low TiO<sub>2</sub> (<1 wt. %)  
313and Zr (<64 ppm) values similar to komatiites (Table A3, Supplementary material 5) and  
314occur at the extension of chemical trends of komatiites (Fig. 6). Therefore, the protolith of  
315group–II metavolcanic rocks was komatiitic basalt.

316 The group–III rocks have higher Fe<sub>2</sub>O<sub>3</sub><sup>T</sup> and lower MgO concentrations (Mg#<40)  
317than group–II komatiitic basalts (Table 3). These also contain higher concentrations of Al<sub>2</sub>O<sub>3</sub>,  
318HFSEs (TiO<sub>2</sub>~0.86-1.91 wt. %; Zr~84–239 ppm and Nb~4.6–12.4 ppm) and REEs relative to  
319the first two groups (Table A3, Supplementary material 5) and plot at the high Fe tholeiitic  
320basalt zone in Mg–Fe+Ti–Al triangular diagram (Fig. 5C; after Jensen, 1976).

321 The trace element<sub>CI</sub> patterns (CI chondrite normalized) of komatiites, komatiitic  
322basalts and tholeiitic basalts in multi–element spider diagram (Fig. A11, supplementary  
323material 3) shows that all elements are enriched (>1) relative to the CI chondrite values. The

324 patterns are characterized by overall enrichment of LREE<sub>CI</sub> relative to HREE<sub>CI</sub> and Ba, Rb,  
 325 Th and U relative to other elements. While most komatiites have depleted HREE<sub>CI</sub> (Gd/Yb<sub>CI</sub>  
 326 ~ 1.38–2.68) relative to LREE<sub>CI</sub> and MREE<sub>CI</sub>, some komatiites (E24 and E35) are also  
 327 enriched in HREE<sub>CI</sub> (Gd/Yb<sub>CI</sub> ~ 0.453–0.632) (Table 3). Komatiitic basalts are overall  
 328 enriched in LREE<sub>CI</sub> and MREE<sub>CI</sub> (La/Yb<sub>CI</sub> ~ 1.09–4.04; Gd/Yb<sub>CI</sub> ~ 1.06–1.97) relative to  
 329 HREE<sub>CI</sub> (exception for sample E34 and A10 having depleted LREE<sub>CI</sub>; La/Yb<sub>CI</sub> ~ 0.669–  
 330 0.784). Tholeiitic basalts have also enriched LREE<sub>CI</sub> and MREE<sub>CI</sub> (La/Yb<sub>CI</sub> ~ 2.30–5.98; Gd/  
 331 Yb<sub>CI</sub> ~ 1.19–1.60) relative to HREE<sub>CI</sub> (Table 3).

### 3327. Discussion

#### 3337.1. Basement and age of the BGB rocks

334 The consisting minerals of metaconglomerate clasts have retained their primary  
 335 igneous textures even after multiphase deformation and metamorphism. The mineralogy  
 336 (quartz with minor K-feldspar, apatite and zircon) and igneous texture of the clasts indicate  
 337 that they were derived from felsic plutonic igneous rocks dated before the BGB rocks. The  
 338 analysed zircon grains present in the clasts show igneous oscillatory growth zoning (Fig. 2E-  
 339 2L) as well as U/Th ratios between 0.33 and 1.56 (Table 2) indicating their igneous origin  
 340 (Mass et al., 1992). The U–Pb ages of the zircons range from 3.51 Ga. to 3.25 Ga. with two  
 341 major peaks around 3.45 Ga. and 3.30 Ga. indicating two major source of the zircons. The  
 342 older peak matches well with the OMTG rocks and the younger peak matches with the  
 343 OMTG granite and third phase of SBG Type–A (Table 1). This implies that the TTGs and  
 344 granites of OMTG and granites of third phase SBG Type–A were the exposed rocks during  
 345 the deposition of BGB sediments and volcanics. They acted as the basement for BGB,  
 346 supplied sediments and possibly contaminated the volcanic rocks. The youngest zircon of the  
 347 suite is aged 3.25 Ga. which is the maximum depositional age of the BGB metaconglomerate.

348The metaconglomerate horizon is one of the basal unit of BGB stratigraphy (Fig. A5c,  
349supplementary material 1). This implies that the older age limit of the BGB rocks is 3.25 Ga.  
350which is 50 Ma. younger than the previously reported older age limit by Saha (1994). Hence,  
351the more precise geological time span of BGB rocks is between 3.25 Ga. and 3.1 Ga.

## 3527.2. Petrogenesis of the BGB metavolcanic rocks

353 The metamorphosed volcanic rocks of the BGB are mostly massive to  
354amygdular/vesicular and occasionally have intact pillow structures (Fig. A2, supplementary  
355material 1). The presence of pillows as well as intercalated sedimentary rocks (e.g. BIF)  
356clearly indicate a submarine depositional setting. Metamorphism and deformation have  
357obscured original mineralogy and textures. However, they can be classified into komatiites,  
358komatiitic basalts and tholeiitic basalts based on their geochemical characteristics.

359 There are two types of komatiites present in the BGB, Al-depleted (D84, E32, E25,  
360D118and D120 having  $\text{Al}_2\text{O}_3/\text{TiO}_2$  ratios between 2.43 and 18.54) and Al-undepleted (E35  
361and E24 having  $\text{Al}_2\text{O}_3/\text{TiO}_2$  ratios around 61.6). The Al-depleted komatiites have depleted  
362 $\text{HREE}_{\text{Cl}}$  ( $\text{Gd}/\text{Yb}_{\text{Cl}} \sim 1.20\text{--}2.68$ ), whereas, Al-undepleted komatiites have enriched  $\text{HREE}_{\text{Cl}}$   
363( $\text{Gd}/\text{Yb}_{\text{Cl}} \sim 0.45\text{--}0.63$ ; Table 3). When melt is produced by moderate degree mantle melting  
364at a higher depth (>9 GPa.), ferromagnesian minerals like olivine and orthopyroxene are  
365melted while garnet being stable at high pressure remains in residual solid. Melting in this  
366way is responsible for the formation of Al- and  $\text{HREE}_{\text{Cl}}$ - depleted komatiites. Besides, when  
367higher degree melting of mantle takes place at relatively shallower depth (<7 GPa.), garnet is  
368melted along with olivine and orthopyroxene, and the magma is enriched in Mg as well as  
369undepleted in Al and  $\text{HREE}_{\text{Cl}}$  (Green, 1975; Sun and Nesbitt, 1978; Herzberg and Ohtani,  
3701988; Hetzberg, 1992).



371 The intercalated volcanic flows of different compositions occur all over the  
372 stratigraphic columns of the BGB without any distinct consecutive order (Supplementary  
373 material 1). For example, komatiite (sample D120) and komatiitic basalt (sample D123)  
374 flows occur at upper stratigraphic position of tholeiitic basalt (sample D33) flows in section  
375 AB (Fig. A1, supplementary material 1), whereas, the basal part of the section is occupied by  
376 pillows of komatiitic basalt (sample A10). The basal volcanic part of section KM is  
377 dominated by tholeiitic basalt flows (Fig. A3, supplementary material 1), whereas, basal part  
378 of section NP is majorly occupied by komatiitic basalt flows (Fig. A5, supplementary  
379 material 1). This type of situation is not uncommon in other Archean greenstone successions.  
380 For example, Cattell (1987) reported intercalated sequences of komatiites and komatiitic  
381 basalts bounded by tholeiitic basalts in some parts of Abitibi greenstone belt, Canada. Such  
382 episodic change of erupting magma composition may indicate two major possibilities, 1)  
383 generation of different magmas from different sources and 2) different degree of partial  
384 melting of a source accompanied by variable degree of fractional crystallization, assimilation  
385 and mixing.

386 The studied rocks form continuous data arrays in different binary geochemical plots  
387 of elements such as compatible vs. incompatible, compatible vs. compatible (Fig. 6) and  
388 incompatible vs. incompatible (Fig. 7 and Fig. A10, supplementary material 3). The  
389 komatiitic basalts followed by tholeiitic basalts plot at the extension of chemical trends  
390 defined by komatiites. The continuous projection of the rock compositions in binary diagrams  
391 indicate an apparent genetic link of the corresponding melts and eliminate the possibility of  
392 different magmatic sources for these rocks. The incompatible (Al, V, Zr, Th, Sm, Nb, Y and  
393 Yb)- and compatible (Mg, Cr and Ni)- elements exhibit respective rise and fall in abundance  
394 along with the melt evolution from komatiitic to tholeiitic. Interestingly, the compositions  
395 mostly follow curvilinear paths (projection of best fit polynomial trend) in the binary

396 diagrams (Fig. 6 and 7; Fig. A10 in supplementary material 3). The slope of most trend lines  
397 becomes steeper (for MgO vs. incompatible elements, and Sm vs. Th) or gentler (for MgO vs.  
398 Ni, Zr vs. Y and Yb, and Th vs. Yb) for komatiitic- and tholeiitic- basalts. During the  
399 formation of basalts, the concentration changing rate of some elements in melt was higher  
400 relative to the others (e.g., Ni, V, Al and Zr relative to MgO; Zr relative to Y and Yb, and Th  
401 relative to Sm and Yb), and this resulted such curvilinear projections of trend lines. The trend  
402 line slope for komatiites is negative in MgO vs. CaO and Sc diagrams and it turns positive for  
403 basalts. This indicates different nature of concentration changing of these elements during  
404 komatiites and basalts formation. Such change of elemental ratios in binary diagrams can be  
405 correlated with the melt evolution processes including fractional crystallization and crustal  
406 contamination in the following section.

#### 407 7.2.1. *Role of fractional crystallization and crustal contamination in melt chemistry*

408       The positive correlation of Cr and Ni with MgO defines simultaneous elimination of  
409 MgO, Cr and Ni from initial melt by fractional crystallization. The chemical trend of MgO  
410 with Cr is linear, but the trend slope with Ni becomes gentler for komatiitic- and tholeiitic-  
411 basalts (Fig. 6). The MgO rich early crystallizing minerals, usually, olivine and  
412 orthopyroxene have high partitioning coefficient values for Ni which result in rapid depletion  
413 of Ni in melt (Arndt et al., 2008). However, Cr does not partition into olivine, rather positive  
414 correlation of Cr with MgO may only be generated by chromite crystallization alongside  
415 olivine (Barnes, 2000). The relationship of CaO and MgO partitioning from melt was  
416 negative during komatiite formation and positive during komatiitic- and tholeiitic- basalt  
417 formation. This indicates that CaO was not removed from parental melt during the formation  
418 of komatiites and minor/no crystallization of CaO rich minerals occurred alongside MgO rich  
419 minerals. However, during the formation of komatiitic- and tholeiitic- basalt, CaO was  
420 partitioned from original melt together with MgO. This may define fractional crystallization

421 controlled by CaO–MgO rich minerals (usually clinopyroxene) and CaO rich minerals  
422 (usually plagioclase) along with MgO rich minerals (olivine and orthopyroxene). But, unlike  
423 CaO,  $\text{Al}_2\text{O}_3$  shows negative trend with MgO for komatiitic– and tholeiitic– basalts which  
424 rules out the possibility of plagioclase fractionation and its control on melt compositions. The  
425 fractionation behaviour of Sc (which is highly compatible in clinopyroxene relative to others)  
426 similar to CaO also supports this event. The negative correlation of  $\text{SiO}_2$ ,  $\text{Al}_2\text{O}_3$ , V, Zr, Nb  
427 and Yb with MgO indicates progressive increase of  $\text{SiO}_2$ ,  $\text{Al}_2\text{O}_3$ , V, Zr and Yb values and  
428 depletion of MgO in melt due to incompatibility of these elements in early crystallizing MgO  
429 rich minerals e.g olivine and pyroxenes. The trend slopes for HFSEs become steeper for  
430 komatiitic– and tholeiitic– basalts indicating rapid increase of those elements in melt with the  
431 decrease of MgO concentrations. This is apparently unexpected as the partitioning coefficient  
432 values of HFSEs are higher (though not significantly) in clinopyroxene relative to olivine and  
433 orthopyroxene (GERM database) which should have caused an opposite situation. This may  
434 indicate either rapid crystal growth of minerals in komatiites leading to disequilibrium trace  
435 elemental partitioning or increase of HFSEs concentration in evolved melts by crustal  
436 contamination accompanied by fractional crystallization. However, significant disequilibrium  
437 trace element partitioning in minerals of komatiites would cause scattered plot of the trace  
438 elements which is mostly absent (exception for Cr) in the binary diagrams (Fig. 6).

439 By definition, incompatible elements (having low partitioning coefficients for early  
440 crystallizing minerals, GERM database) do not participate significantly in fractional  
441 crystallization processes. Their abundance in melt though increases with fractional  
442 crystallization, the elemental ratios remain unaffected. Therefore, mantle generated melt that  
443 is not contaminated with crustal components would exhibit the incompatible elemental  
444 patterns equivalent to that of mantle. Zirconium and Th are highly enriched in continental  
445 crust relative to Sm, Y and Yb (Taylor and McLennan, 1985). Contamination with

446continental crust would result in increase of Th/Yb, Th/Sm, Zr/Y and Zr/Yb ratios. Most  
447studied rock samples exhibit the said ratios higher than mantle ratios (Fig. 7) indicating their  
448contamination with crustal components. Such assimilation with crustal components might  
449had happened directly, by thermal erosion of crustal rocks within crustal chambers, magma  
450tube and at surface, or indirectly, by mixing of melt with crustal component rich continental  
451lithospheric mantle, or all of the processes together. To understand the crust vs. mantle  
452contamination process, we plotted the studied rock compositions in La/Sm vs. La/Ta binary  
453diagram (Fig. 8) after Lassiter and DePaolo (1997). Most of the komatiites and basalts of the  
454BGB follow trend parallel to crustal contamination line in the diagram and therefore indicate  
455the possibility of their direct assimilation with crustal material.

#### 4567.2.2. *Modelling of melting, fractional crystallization and crustal contamination*

457 To understand the melting and assimilation–fractional crystallization (AFC) process  
458responsible for different melt composition in the studied area, we used “PETROMODELER”  
459excel spreadsheet (Ersoy, 2013). We have tried to address the possible ways of changing  
460magmatic compositions using Zr/Y vs. Yb AFC modelling (Fig. 9). For melting modelling,  
461we used primitive mantle (PM) data of Palme and O’Neill (2014) as melt source composition  
462( $C_o^M$ ). For assimilation–fractional crystallization modelling, we have used data of  
463granodiorite (3.3 Ga; Nelson et al., 2014) occurring nearby the BGB as a possible  
464contaminant ( $C_a$ ). Partitioning coefficient data of Yb, Zr and Y used for fractional  
465crystallization modelling are taken from GERM database. The results of the melting and AFC  
466modelling are as follows.

467 The Al–depleted komatiite compositions could be generated by 25% garnet–lherzolite  
468facies melting (Walter, 1998) of PM and followed by AFC~ 18–36% and r (ratio of  
469assimilation rate and fractional crystallization rate) ~ 0.27 (Fig. 9A). The Al–undepleted

470 komatiite compositions (e.g. E35 and E24) could be formed by different degree spinel–  
471 herzolite facies melting (Kinzler, 1997) of PM. The E35 composition was possibly formed  
472 by 29% melting followed by AFC ~ 18% with  $r = 0.27$ , whereas, E24 was formed by 23%  
473 melting followed by AFC ~ 12% with  $r = 0.27$  (Fig. 9B). The komatiitic basalts could not be  
474 generated from the Al depleted komatiites or Al–undepleted E35 by AFC. The melt  
475 compositions generated from these komatiites do not reach the komatiitic basalt compositions  
476 successfully and only the E24 melt composition is suitable for being the parental melt  
477 composition. The reason is that E24 has the lowest incompatible element contents as well as  
478 it exhibits incompatible elemental ratios same as mantle (Fig. 7). Eventually, Al–undepleted  
479 E24 komatiite was the least fractionated and contaminated melt (most primitive) composition  
480 among other komatiites. Therefore, the komatiitic basalts could be the products of Al–  
481 undepleted E24 komatiite melt or similar melt compositions formed by moderate degree  
482 melting of mantle.

483        Most komatiitic basalts (exception for E38 and E39) could be generated by AFC of  
484 E24, where AFC varied from 18 to 50 % and the  $r$  value was between 0.10 and 0.28 (Fig.  
485 9C). The more evolved komatiitic basalt compositions, like E38 and E39 could be reached by  
486 AFC of other more primitive komatiitic basalt compositions such as E34. The E38 and E39  
487 compositions could be formed by AFC ~30% and  $r = 0.23$  of E34. The tholeiitic basalts were  
488 generated by AFC of various komatiitic basalt compositions. For example, the D33 melt  
489 composition was the result of AFC ~18% with  $r = 0.27$  of D123. The D111 was formed by  
490 AFC ~ 50% with  $r = 0.27$  of E38. The E17, E13 and E11 was formed by 54% fractional  
491 crystallization without significant assimilation of E38 melt composition. The most evolved  
492 tholeiitic basalt E22 could be formed by AFC ~30% with  $r = 0.59$  of E17 or similar melt  
493 compositions (Fig. 9D).

494 To verify the AFC model, we tested the melt composition evolution on multi-element  
495(HFSEs and REEs) spider-diagram (Fig. 10). We selected some melt compositions those  
496represents each melt groups e.g., Al-undepleted komatiite (E24, starting composition),  
497primitive komatiitic basalt (E34), evolved komatiitic basalt (E38) and tholeiitic basalt  
498(D111). The E34 composition could be successfully reached by AFC ~ 45% ( $r=0.14$ ) of E24  
499similar to the previous modelling results. The composition E34 could be evolved to E38 by  
500AFC~ 36% ( $r=0.35$ ) and composition E38 could be evolved to D111 by AFC~ 45% ( $r=0.27$ ).  
501The multi-element AFC results match well with the previously determined bi-variant AFC  
502results.

503 In summary (Fig.11), the more primitive melt compositions (komatiitic) were  
504episodically introduced into the magma chambers of the studied greenstone belt and were  
505evolved by fractional crystallization and assimilation. The komatiite melts were generated at  
506different depth of mantle with different melting degrees, possibly from rapidly ascending  
507mantle plume. As stated in the introduction section, the genetic relationship of komatiites,  
508komatiitic basalts and tholeiitic basalts are highly debated instead of their close field  
509association. Nesbitt and Sun (1976) and Arndt and Nesbitt (1984) proposed that the basalts  
510were formed by low–moderate degree melting of different parts of the mantle source that  
511produced komatiites. Campbell et al. (1989) proposed that komatiites were originated from  
512hot axis of mantle plume and the other basalts are melting product of cooler plume annulus.  
513However, our study suggests that the komatiitic basalts could be generated by AFC of Al–  
514undepleted komatiites (e.g. E24) formed by moderate degree melting at relatively lower  
515pressure. It was also noted that the tholeiitic basalts could not be formed by even 90% AFC  
516of a komatiite melt, but, they could be formed by AFC of evolved komatiitic basalt melts.  
517However, it is difficult to imagine that a melt composition (e.g. E38) generated by 30% AFC,  
518again participated in 50–54% AFC to form another melt (e.g. D111 or E17). This is possible

519only if the residual melt of evolved komatiitic basalt composition was separated out in a  
520different magma chamber and went through another AFC process.

521       The komatiite-basalt succession of the BGB, Singhbhum Craton may be geologically  
522compared with similar type Archean volcanic successions present in Stoughton-Rouemature  
523group (~2.7 Ga), Abitibi Greenstone belt, Canada (Dostal and Mueller, 2013),  
524Koolyanobbing greenstone belt (~3.0 Ga), Yilgarn Craton, Australia (Angerer et al., 2013),  
525and Onverwacht suite (~3.5-3.3 Ga), Barberton greenstone belt, South Africa (Furnes et al.,  
5262012). The field association, metamorphic grade, mineralogy and major elemental  
527compositions of the komatiite-basalt suites in the BGB and the mentioned three Archean  
528greenstone belts are very alike. While the Abitibi- and Koolyanobbing- greenstone belt  
529contains both Al- depleted and -undepleted variety of komatiites along with other basalts, the  
530Barberton greenstone belt contain mainly Al-undepleted variety of komatiite associated with  
531basalts. Dostal and Mueller (2013) divided the volcanic assemblage of Stoughton-  
532Rouemature group into two sets based on Al- and HREE- undepletion/depletion. They  
533suggested that the different komatiites and basalts were generated by melting of different  
534parts of a heterogeneous mantle plume and denied the possibility of basalt formation from  
535komatiites by AFC processes. Similarly Angerer et al. (2013) suggested that the different  
536volcanic rocks of Koolyanobbing greenstone belt were generated from a zoned plume and/or  
537melting over a range of depths in a convergent plate margin setting. Furnes et al. (2012) also  
538concluded that the komatiites and basalts of Onverwacht suite were generated by variable  
539degrees of partial melting at different depths and temperatures of metasomatised mantle in a  
540subduction zone setting.

541       When we plotted the data of komatiites and basalts present in these greenstone belts  
542together in binary diagrams such as MgO-CaO, MgO-Ni, MgO-Zr and Zr-Yb (Fig.12), the  
543basalts and komatiites formed continuous data array. They also showed

544olivine/orthopyroxene dominated fractional crystallization during the formation of komatiites  
545(Fig. 12A), clinopyroxene dominated fractional crystallization during the formation of basalts  
546(Fig. 12B), and incorporation of crustal components during the formation of basaltic melts  
547(Fig. 12C and D) similar to the BGB rocks. Therefore, we propose that the different  
548komatiites of these greenstone belts were possibly generated from different parts of  
549heterogeneous sources, but the basalts were possibly evolved from the corresponding  
550komatiites involving AFC processes similar to the BGB rocks. Most of the studies (e.g.  
551Nesbitt and Sun, 1976; Arndt and Nesbitt, 1984; Arndt et al., 1997; Hanski et al., 2001;  
552Furnes et al., 2012; Dostal and Mueller, 2013; Angerer et al., 2013; Verma et al., 2017)  
553though suggest diverse origins of Archean komatiites and basalts, the possibility of a genetic  
554connection among them cannot be completely ruled out. Therefore, a geochemical  
555reinvestigation is needed to verify the rationality of the AFC model as an effective  
556mechanism for basalt formation from komatiites in Archean greenstone belts.

## 5578. Conclusions

558       The results and discussion led us to the following major conclusions. The melt of Al–  
559depleted and –undepleted komatiites present in the BGB were originated by moderate–high  
560degree melting of mantle plume at different depth. The melts further went through  
561assimilation and olivine–orthopyroxene controlled fractional crystallization in magma  
562chambers. The komatiitic basalts were mostly produced from clinopyroxene controlled  
563fractional crystallization and assimilation of Al–undepleted komatiite melts formed by  
564moderate degree melting of mantle plume. The tholeiitic basalts were generated by mostly  
565clinopyroxene controlled fractional crystallization and assimilation of more evolved  
566komatiitic basalt and primitive tholeiitic basalt melts. The Older Metamorphic Tonalitic  
567Gneiss and third phase of Singhbhum Granite Type–A were the basement of the Badampahar  
568greenstone belt (BGB) rocks and they contributed sediments for metaconglomerate and



569contaminated the magma. The newly discovered older age limit of the BGB rocks is 3.25 Ga  
570which is 50 Ma younger than the previously reported age limit. The formation of the BGB  
571rocks is confined between 3.25 Ga and 3.1 Ga.

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762**Table 1.** Early to middle Archean stratigraphic sequence of rocks in the Singhbhum Craton, modified  
763after Saha (1994).

Stratigraphic unit	Events	Major lithologies	Age (Ga)	References
	<i>Metamorphism of OMG, OMTG and SBG</i>		3.02-2.96, 2.52 and 1.06	Upadhyay et al. (2014)
<b>Mayurbhanj Granite</b>		Granite	3.09-3.08	Mishra et al. (1999)
<b>SBG Type-B</b>	Emplacement of granitic pluton	Granodiorite to granite	3.12	Ghosh et al. (1996)
<b>Bonai granite</b>		Granite to granodiorite	3.16	Sengupta et al. (1991)
	<i>Metamorphism of SBG Type-A, OMTG and OMG</i>		3.19 - 3.13	Upadhyay et al. (2014)
<b>Iron Ore Group</b>	Deposition and metamorphism of Iron ore group of rocks	Mafic to felsic volcanic rocks, tuff, banded iron formations, local dolomite, quartzitic sandstone and conglomerate	3.3-3.1	Saha (1994)
-----Unconformity-----				
	<i>Metamorphism of SBG Type-A, OMTG and OMG</i>		3.24 3.34 - 3.26	Mishra et al. (1999) Upadhyay et al. (2014)
<b>SBG Type-A:</b>				
Third phase			3.33 3.32	Upadhyay et al. (2014) Mishra et al. (1999)
Second phase	Emplacement of granitic pluton	Granite, tonalite and granodiorite	3.35	Upadhyay et al. (2014)
First phase			3.44	Upadhyay et al. (2014)
	<i>Metamorphism of OMG and OMTG</i>		3.40	Mishra et al. (1999)
<b>OMTG granite</b>		Granite	3.32	Upadhyay et al. (2014)
<b>OMTG</b>	Intrusion of tonalite-trondhjemite-granodiorite rocks in OMG	Tonalite gneiss and granodiorite	3.44 and 3.52	Acharyya et al. (2010)
			3.45 - 3.43	Mishra et al. (1999)
			3.45 - 3.44	Upadhyay et al. (2014)
<b>OMG</b>	Deposition of sediments with mafic volcanism	Amphibolites, pelitic schists, banded calc-gneiss	3.55-3.44	Mishra et al. (1999)
-----Unconformity-----				
<b>Unpreserved primitive crust represented by xenocrystic zircons present in younger rocks</b>			3.61 3.6-3.55	Upadhyay et al. (2014) Goswami et al. (1995)

OMG ~ Older Metamorphic Group; OMTG ~ Older Metamorphic Tonalite Gneiss; SBG ~ Singhbhum Granite

764**Table 2.** Summary of U-Pb isotope data and concordant ages of zircons extracted from conglomerate  
765sample from the Badampahar greenstone belt, Singhbhum Craton, Eastern India..

sample	U [ppm]	Th [ppm]	Th/U	Pb207/ U235	Error [sigma a]	Pb206/ U238	Error [sigma]	Age [Ma]	Error [sigma]
<b>A33_G0</b> <b>01</b>	104.13 43	52.014 26	0.4994 92	29.309 01	0.382 89	0.7176 9	0.00869	3450 .8	18.35
<b>A33_G0</b> <b>02</b>	47.045 45	33.602 12	0.7142 48	24.863 15	0.343 16	0.6681 1	0.0085	3305 .9	19.65
<b>A33_G0</b> <b>03</b>	186.06 74	79.705 95	0.4283 71	30.463 04	0.391 38	0.7263 9	0.00863	3492	18.01
<b>A33_G0</b> <b>05</b>	263.24 31	111.38 27	0.4231 17	29.407 19	0.376 92	0.7050 4	0.00833	3483 .5	18.01
<b>A33_G0</b> <b>06</b>	66.603 67	56.897 96	0.8542 77	28.881 05	0.390 74	0.6963 6	0.00869	3474 .8	19.02
<b>A33_G0</b> <b>07</b>	394.86 46	132.04 72	0.3344 11	23.990 5	0.311 14	0.6516 6	0.00776	3289	18.54
<b>A33_G0</b> <b>08</b>	45.459 65	32.862 08	0.7228 85	24.744 61	0.350 25	0.6669 7	0.00863	3301 .1	20.23
<b>A33_G0</b> <b>09</b>	283.32 99	204.10 78	0.7203 89	30.666 53	0.397 15	0.7353 6	0.00871	3483 .3	18.26
<b>A33_G0</b> <b>10</b>	164.39 48	93.635 89	0.5695 79	25.438 96	0.334 69	0.6837 3	0.0082	3305 .5	18.83
<b>A33_G0</b> <b>11</b>	135.32 17	93.530 97	0.6911 75	25.431 39	0.337 74	0.6830 5	0.00825	3306 .6	19.02
<b>A33_G0</b> <b>12</b>	468.86 87	483.36 99	1.0309 28	26.544 42	0.345 11	0.6574	0.00775	3433 .2	18.48
<b>A33_G0</b> <b>14</b>	56.031 66	70.024 96	1.2497 39	26.723 57	0.374 03	0.7046 2	0.0089	3335 .6	20.03
<b>A33_G0</b> <b>15</b>	190.82 48	200.64 97	1.0514 86	28.579 59	0.380 9	0.6894 5	0.00826	3473 .9	18.94
<b>A33_G0</b> <b>16</b>	100.43 41	71.913 05	0.7160 22	25.491 3	0.348 73	0.6856 8	0.0084	3304 .3	19.71
<b>A33_G0</b> <b>19</b>	168.09 5	87.760 32	0.5220 88	29.871 07	0.407 72	0.7242 4	0.00876	3466 .1	19.5
<b>A33_G0</b> <b>20</b>	238.92 74	364.64 39	1.5261 7	22.868 83	0.314 22	0.6260 5	0.00757	3276 .7	19.98
<b>A33_G0</b> <b>21</b>	367.37 74	161.37 77	0.4392 7	23.937 66	0.326 32	0.6666 1	0.00796	3249 .8	19.89
<b>A33_G0</b> <b>27</b>	269.58 63	260.74 81	0.9672 16	26.205 56	0.372 09	0.6646	0.00804	3396 .3	20.67
<b>A33_G0</b> <b>28</b>	205.09 7	212.44 39	1.0358 22	27.490 39	0.394 88	0.6887 3	0.0084	3415 .3	20.9
<b>A33_G0</b> <b>32</b>	213.02 6	145.71 4	0.6840 2	28.546 2	0.404 31	0.6948 6	0.0082	3460	21.1
<b>A33_G0</b> <b>33</b>	178.66 7	86.546 71	0.4844 02	30.328 9	0.434 35	0.7263 4	0.00867	3485 .2	21.27
<b>A33_G0</b> <b>34</b>	200.86 82	240.16 18	1.1956 19	25.231 24	0.361 53	0.6294 5	0.00748	3421 .9	21.46
<b>A33_G0</b> <b>35</b>	169.15 22	204.84 02	1.2109 82	32.381 92	0.517 54	0.7754 9	0.01055	3485 .2	23.28
<b>A33_G0</b> <b>36</b>	191.35 34	127.84 54	0.6681 11	30.408 4	0.436 9	0.7371 2	0.00874	3466 .4	21.44
<b>A33_G0</b>	264.30	236.12	0.8934	29.924	0.427	0.7177	0.00839	3482	21.39

<b>38</b>	03	97	14	45	68	5		.8	
<b>A33_G0</b>	201.39	195.67	0.9716	24.850	0.362	0.668	0.00792	3305	22.13
<b>40</b>	68	96	12	14	07			.3	
<b>A33_G0</b>	219.36	95.851	0.4369	30.604	0.446	0.7344	0.00867	3481	21.88
<b>42</b>	92	15	4	99	28	8		.9	
<b>A33_G0</b>	84.047	33.181	0.3948	30.450	0.465	0.7296	0.00913	3484	22.78
<b>43</b>	49	92		62	88	2		.4	
<b>A33_G0</b>	294.43	297.22	1.0094	27.682	0.403	0.6767	0.00792	3453	22.03
<b>44</b>	05	25	83	08	93	2		.3	
<b>A33_G0</b>	186.59	140.30	0.7519	28.012	0.439	0.6987	0.00896	3422	23.5
<b>45</b>	6	23	04	75	88	1		.1	
<b>A33_G0</b>	304.47	207.82	0.6825	23.464	0.346	0.6468	0.00757	3265	22.72
<b>47</b>	39	99	87	46	86	8		.6	
<b>A33_G0</b>	278.04	433.03	1.5574	25.056	0.371	0.619	0.00725	3437	22.53
<b>48</b>	39	43	32	68	6				
<b>A33_G0</b>	192.93	165.88	0.8597	27.113	0.408	0.6560	0.00778	3469	22.82
<b>50</b>	92	04	55	44	28	5		.2	
<b>A33_G0</b>	213.55	262.05	1.2270	24.908	0.379	0.6744	0.00802	3293	23.42
<b>52</b>	46	18	95	94	32	9		.8	
<b>A33_G0</b>	136.90	80.172	0.5855	25.938	0.402	0.703	0.00848	3292	23.84
<b>54</b>	75	41	95	84	17			.4	
<b>A33_G0</b>	40.173	31.832	0.7923	25.200	0.444	0.6829	0.00967	3292	26.67
<b>56</b>	64	69	77	21	69	9		.4	
<b>A33_G0</b>	114.17	61.043	0.5346	30.473	0.487	0.7201	0.00888	3505	24.27
<b>57</b>	77	21	33	58	91	3		.9	
<b>A33_G0</b>	49.688	34.085	0.6859	24.608	0.418	0.6628	0.00874	3302	26.05
<b>59</b>	45	73	89	64	3	3		.2	
<b>A33_G0</b>	274.34	113.63	0.4142	29.118	0.459	0.7009	0.00823	3477	24.25
<b>61</b>	37	53	08	04	37	8		.2	
<b>A33_G0</b>	50.745	23.615	0.4653	29.385	0.505	0.7053	0.00938	3481	26
<b>62</b>	65	73	74	6	47	2		.9	
<b>A33_G0</b>	218.31	119.59	0.5478	23.729	0.381	0.6206	0.00738	3348	24.96
<b>63</b>	2	4	12	12	08	6		.2	
<b>A33_G0</b>	224.12	150.09	0.6696	25.444	0.410	0.6240	0.00743	3448	24.9
<b>64</b>	66	33	81	27	32	1		.5	
<b>A33_G0</b>	156.46	88.595	0.5662	27.977	0.457	0.6719	0.00812	3481	25.15
<b>65</b>	58	02	26	91	52				
<b>A33_G0</b>	250.55	146.83	0.5860	26.028	0.427	0.6324	0.00748	3463	25.49
<b>69</b>	67	72	44	27	44			.1	
<b>A33_G0</b>	66.075	30.460	0.4609	29.087	0.508	0.6906	0.00899	3498	26.73
<b>70</b>	07	45	98	45	77	2		.8	
<b>A33_G0</b>	110.47	47.080	0.4261	29.304	0.499	0.6944	0.00862	3501	26.25
<b>71</b>	75	95	59	18	63			.8	
<b>A33_G0</b>	228.88	321.55	1.4048	27.572	0.462	0.6658	0.00791	3472	26.11
<b>73</b>	4	61	87	75	62	1		.7	
<b>A33_G0</b>	205.09	92.835	0.4526	27.313	0.465	0.6516	0.00778	3491	26.57
<b>76</b>	7	87	44	1	75	7		.3	
<b>A33_G0</b>	75.589	51.032	0.6751	23.458	0.419	0.6336	0.00799	3298	28.18
<b>79</b>	88	24	2	86	53	8		.1	
<b>A33_G0</b>	247.91	122.63	0.4946	25.274	0.439	0.6194	0.00741	3449	27.32
<b>80</b>	37	71	77	31	9	5		.8	
<b>A33_G0</b>	31.187	21.889	0.7018	24.331	0.470	0.6524	0.0092	3309	30.1
<b>81</b>	43	16	58	92	92	2		.7	
<b>A33_G0</b>	26.430	22.787	0.8621	24.135	0.483	0.6476	0.00951	3308	31.06
<b>83</b>	03	24	72	99	21			.7	

766

767

768**Table 3.** Ratios and anomalies of rare earth elements in the BGB metavolcanic rocks and other  
769modern standard volcanic rocks.

	(La/ Sm) <sub>CI</sub>	(Gd/ Yb) <sub>CI</sub>	(La/ Yb) <sub>CI</sub>	(Eu/ Eu*) <sub>CI</sub>	(La/ Sm) <sub>NM</sub>	(Gd/ Yb) <sub>NM</sub>	(La/ Yb) <sub>NM</sub>	(Ce*/ Ce) <sub>CI</sub>	Mg#	Al <sub>2</sub> O <sub>3</sub> / TiO <sub>2</sub>	CaO/ Al <sub>2</sub> O <sub>3</sub>
<b>Komatites</b>	<b>E35</b>	bdl	0.453	bdl	0.988	bdl	0.454	bdl	75.0	61.6	0.369
	<b>D-84</b>	2.28	2.68	9.04	0.939	3.72	2.68	15.4	0.856	56.2	4.20
	<b>E32</b>	2.25	2.20	7.79	0.820	3.67	2.20	13.2	0.764	61.6	4.86
	<b>E25</b>	0.849	1.20	0.854	1.27	1.38	1.20	1.45	0.795	59.9	18.5
	<b>E24</b>	1.25	0.632	0.598	1.04	2.04	0.633	1.02	0.756	65.0	61.6
	<b>D118</b>	1.92	2.08	4.58	0.949	3.12	2.08	7.79	0.978	60.5	7.07
	<b>D120</b>	2.45	1.38	3.59	1.00	3.99	1.38	6.10	0.991	59.7	12.9
<b>Komatitic basalts</b>	<b>E34</b>	0.856	1.06	0.784	0.802	1.40	1.06	1.33	0.964	39.4	27.4
	<b>E39</b>	1.16	1.65	1.86	0.923	1.89	1.65	3.16	0.991	33.1	9.78
	<b>E36</b>	0.788	1.54	1.09	1.03	1.28	1.55	1.86	1.00	45.8	11.6
	<b>E37</b>	0.856	1.52	1.24	1.01	1.39	1.52	2.11	0.994	42.6	10.5
	<b>E40</b>	1.01	1.97	2.15	1.05	1.64	1.98	3.66	0.934	35.1	7.35
	<b>E44</b>	1.05	1.63	1.68	1.02	1.72	1.63	2.85	0.952	40.7	10.6
	<b>E38</b>	1.03	1.64	1.67	1.05	1.06	1.14	1.14	1.03	20.8	10.8
	<b>A-10</b>	0.651	1.14	0.669	0.874	3.64	1.67	6.87	1.03	31.6	9.38
	<b>D11</b>	2.24	1.67	4.04	1.03	3.44	1.23	4.27	0.966	41.0	17.9
	<b>E23</b>	0.956	1.21	0.936	0.994	1.56	1.22	1.59	1.13	50.4	17.6
	<b>D123</b>	1.47	1.60	2.68	0.970	2.40	1.60	4.55	1.03	42.7	13.1
<b>A26</b>	1.25	1.09	1.21	0.938	2.04	1.09	2.05	0.914	45.0	21.7	
<b>Tholeiitic basalts</b>	<b>D33</b>	1.65	1.60	2.98	0.991	2.70	1.60	5.06	0.982	25.1	12.1
	<b>E11</b>	1.95	1.19	2.31	0.908	3.18	1.20	3.92	0.960	24.9	10.7
	<b>E13</b>	1.97	1.21	2.35	0.869	3.21	1.21	4.00	0.977	23.1	10.9
	<b>E17</b>	2.11	1.20	2.51	0.909	1.68	1.64	2.84	0.979	39.4	12.1
	<b>E22</b>	2.13	1.35	3.21	0.863	3.48	1.35	5.45	0.990	19.6	6.81

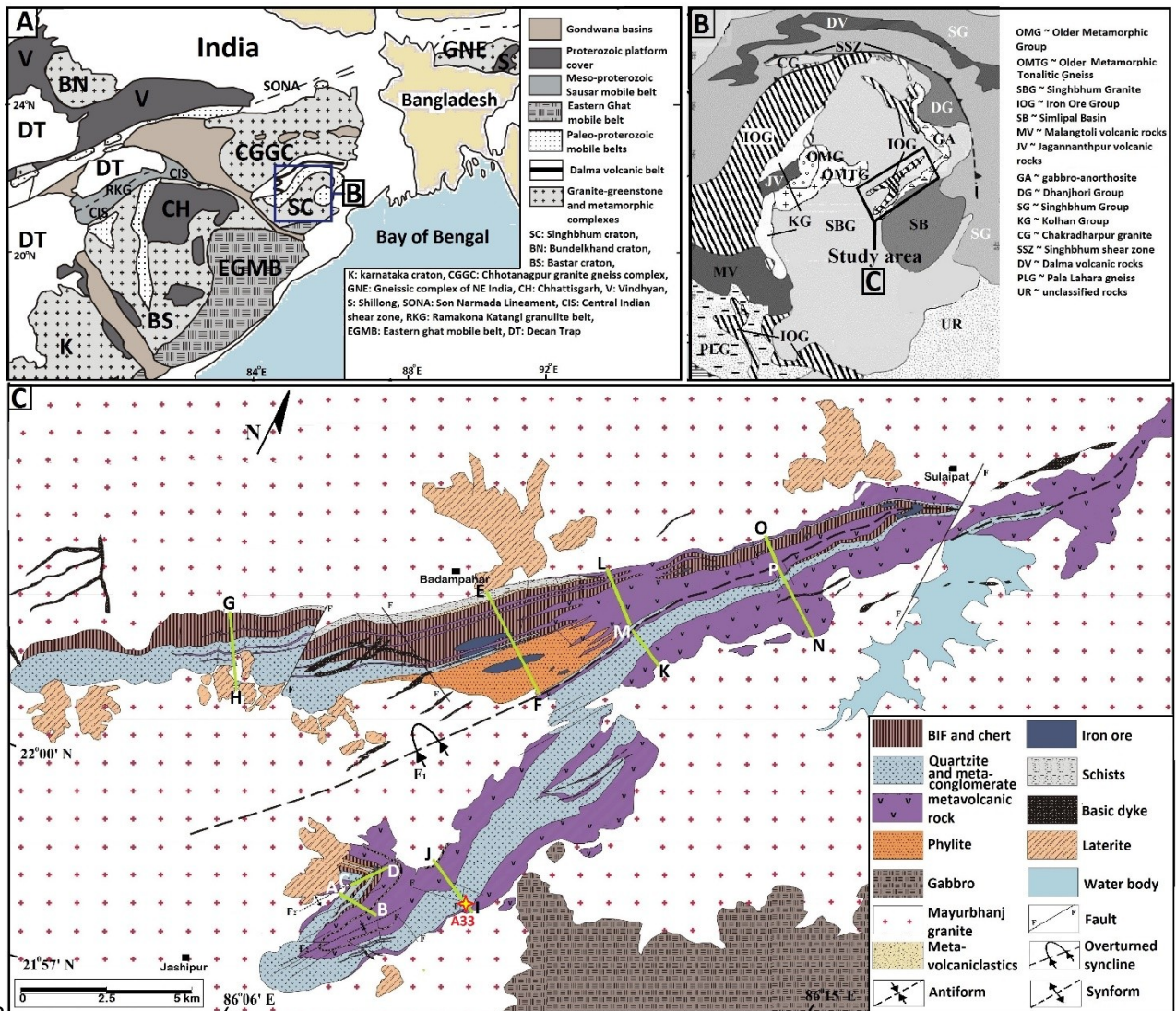
	<b>D111</b>	3.55	1.45	5.98	0.791	5.78	1.45	10.2	0.907	15.4	13.5	0.510
<b>others</b>	<b>PM</b>	1.00	1.00	1.00	0.998	1.63	1.00	1.70	-	-	-	-
	<b>NM</b>	0.614	0.998	0.588	1.00	1.00	1.00	1.00	-	-	-	-
	<b>EM</b>	1.56	1.04	1.91	1.00	2.55	1.04	3.24	-	-	-	-

<sup>CI</sup> ~ CI chondrite normalized; <sup>NM</sup> ~ NMORB normalized; bdl ~ below detection limit; <sup>a</sup> ~ Ghosh and Baidya (2017b); <sup>b</sup> ~ Sun & McDonough (1989); PM ~ average primitive mantle; NM ~ average NMORB; EM ~ average EMORB

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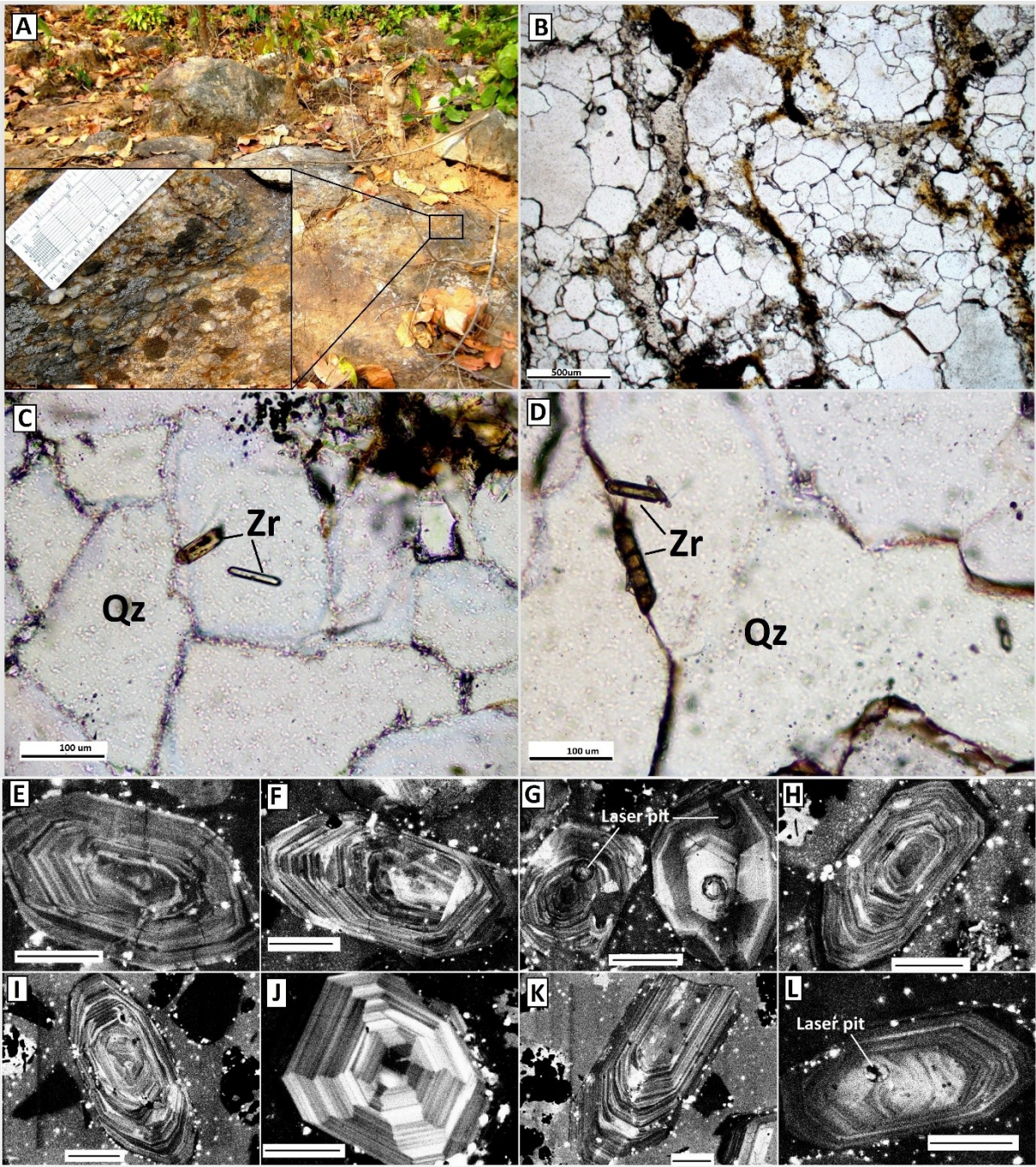
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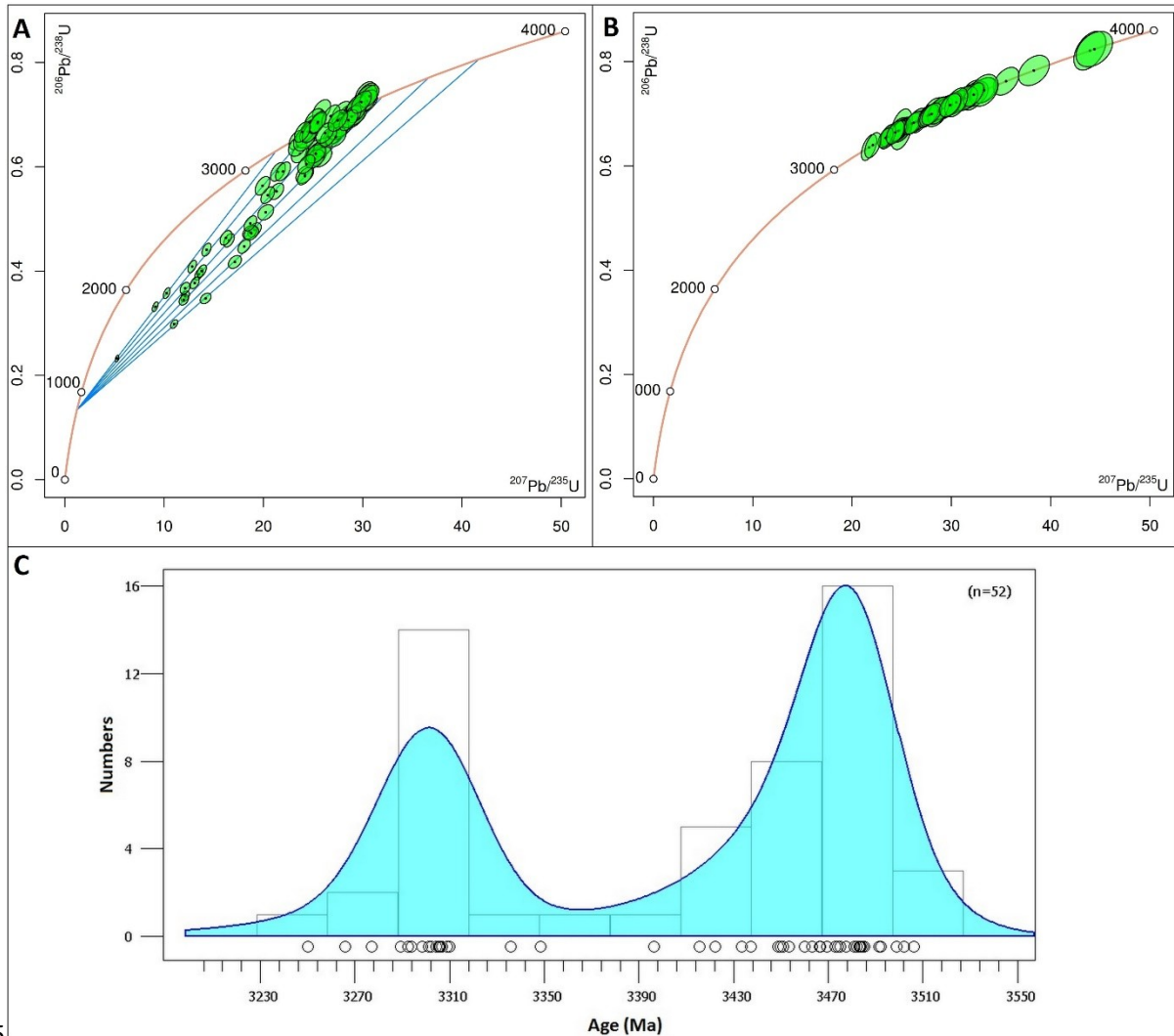
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774**Fig. 1. (A)** Geological map of India showing locations of major Precambrian terrains  
 775(modified after Acharyya, 2002). **(B)** Enlarged view of the Singhbhum Craton with the  
 776spatial distribution of different stratigraphic units (modified after Saha, 1994). **(C)** Geological  
 777map of the Badampahar greenstone Belt, Singhbhum craton (after Ghosh and Baidya, 2017b).  
 778Location ( $N22^{\circ}00'15''$ ,  $E86^{\circ}08'40''$ ) of the dated metaconglomerate sample (A33) is marked.



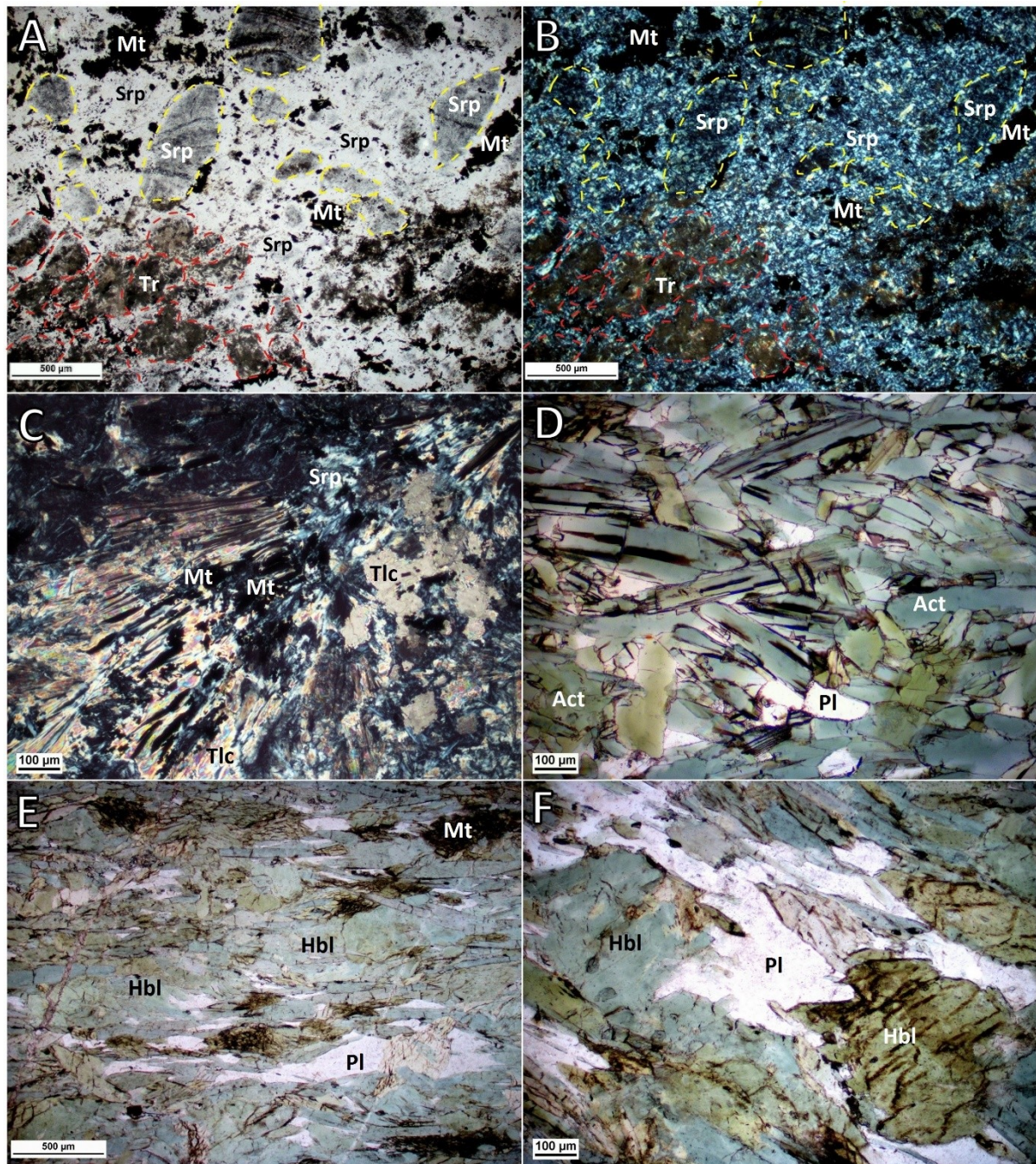


**Fig. 2.** (A) Outcrop of metaconglomerate (A33) at the southern limb of BGB syncline. (B), (C) & (D) Photomicrographs show ellipsoidal clasts of metaconglomerate consisting mostly of quartz crystals. The contact of quartz crystals are mostly sharp and straight. Inclusions of zircon are present within quartz. (E)-(L) CL images of analysed zircon grains showing oscillatory growth zoning.



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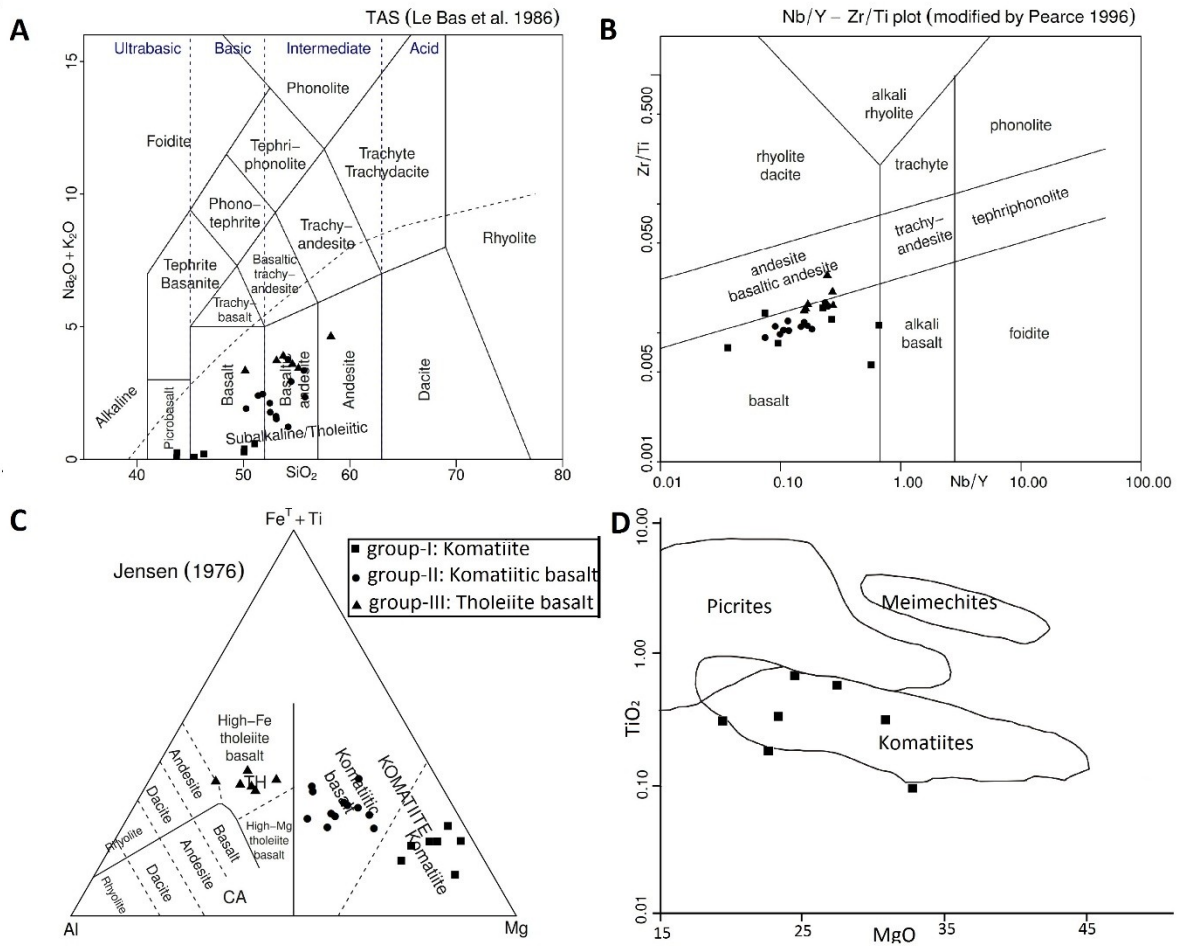
**Fig. 3.** (A) Wetherill concordia diagram of metaconglomerate sample A33, suggesting Archaean protolith ages and a Neoproterozoic partial resetting event. Blue lines mark the trajectory of the Pb-loss. (B) Assuming initial concordance, the original U-Pb compositions may be restored by projecting the data back to the concordia line. (C) Kernel Density Estimate (KDE) of the zircons having concordant ages.



791

792 **Fig. 4.** Photomicrographs of metavolcanic rocks in the Badampahar greenstone belt,  
 793 Singhbhum Craton. **(A)** and **(B)** Serpentine (Srp) and tremolite (Tr) pseudomorphs (marked  
 794 by yellow and red dotted lines, respectively) after original igneous minerals along with  
 795 magnetite (Mt) crystals represent relic cumulate texture in a ‘meta’ komatiite. **(C)** Relic  
 796 spinifex texture in a ‘meta’ komatiite represented by acicular radiating crystals of magnetite  
 797 and Tlc-Srp pseudomorphs after olivine/orthopyroxene. **(D)** Actinolite (Act) and plagioclase  
 798 in a ‘meta’ komatiitic basalt showing crude alignment along regional  $S_1$  foliation. **(E)**  
 799 Hornblende (Hbl) and Pl crystals of a ‘meta’ tholeiitic basalt aligned along the  $S_1$  foliation.

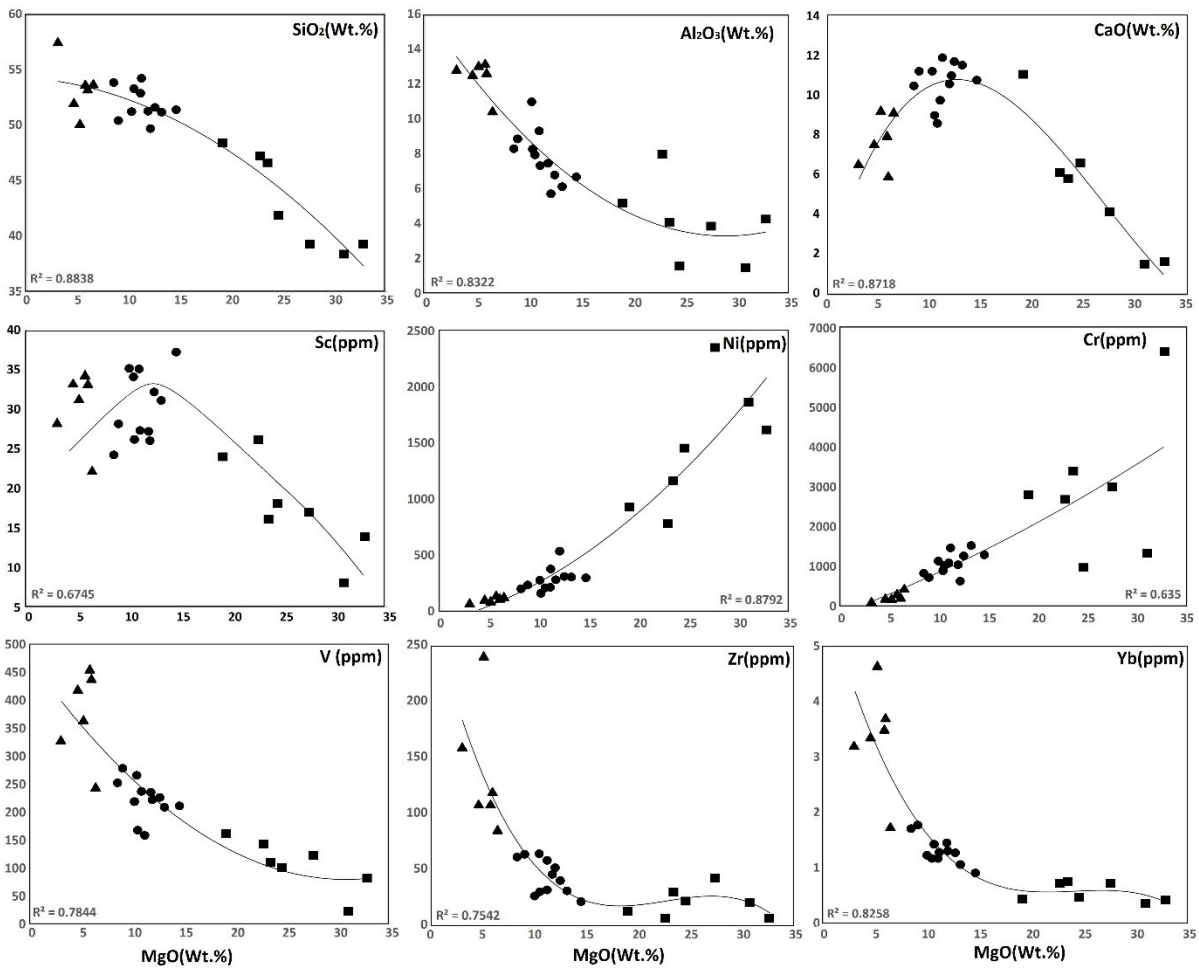
800(F) Relic ophitic and subophitic texture represented by the outline of Hbl and Pl crystals in a 801 'meta' tholeiitic basalt.



802

803 **Fig. 5.** The metavolcanic rocks of BGB are plotted in (A) SiO<sub>2</sub> vs. K<sub>2</sub>O+Na<sub>2</sub>O diagram (after  
 804 Le Bas et al., 1986), (B) Nb/Y vs. Zr/Ti diagram (after Pearce, 1996), (C) Mg-Fe+Ti-Al  
 805 triangular diagram (Jensen, 1976). (D) High Mg group-I rocks of the BGB plotted in MgO-  
 806 TiO<sub>2</sub> diagram (after Arndt et al., 2008).

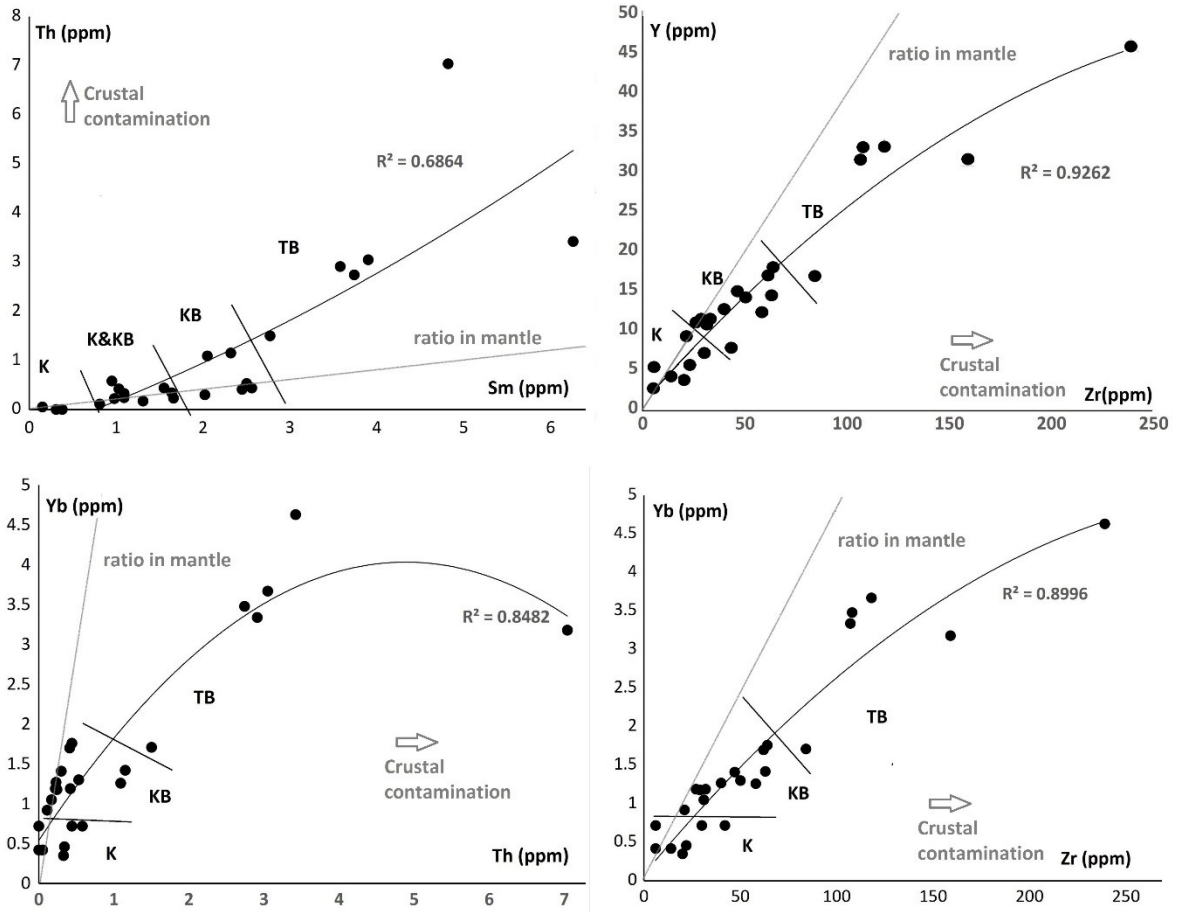
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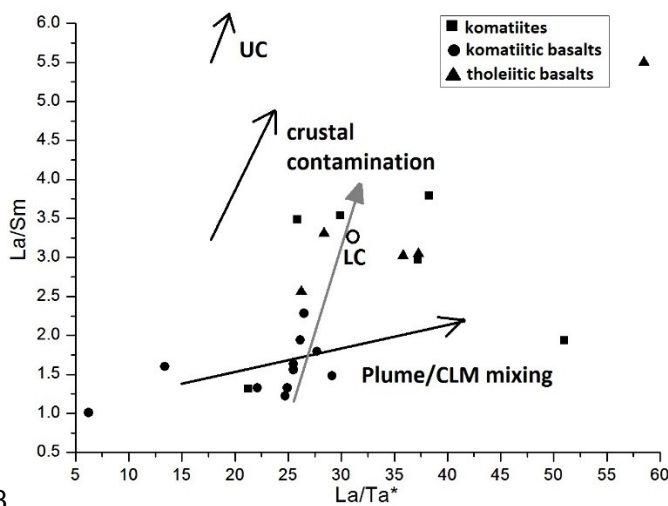
809**Fig. 6.** Bi-variant plots (MgO vs. SiO<sub>2</sub>, Al<sub>2</sub>O<sub>3</sub>, CaO, Sc, Ni, Cr, V, Zr and Yb) of the BGB  
 810metavolcanic rocks are showing continuous data array indicated by the projection of best fit  
 811polynomial trends. The rock type symbols are same as in Fig. 5C.

812



813

814**Fig. 7.** Bi-varient plots (Zr vs. Nb and Sm, Sm vs. Nb and Th vs. Nb) of the BGB  
 815metavolcanic rocks show continuous data array indicated by the projection of best fit  
 816polynomial trends. Grey lines mark the elemental ratios in Earth’s mantle (Plame and  
 817O’Neill, 2014). K~ komatiites, KB~komatiitic basalts, TB~tholeiitic basalts.

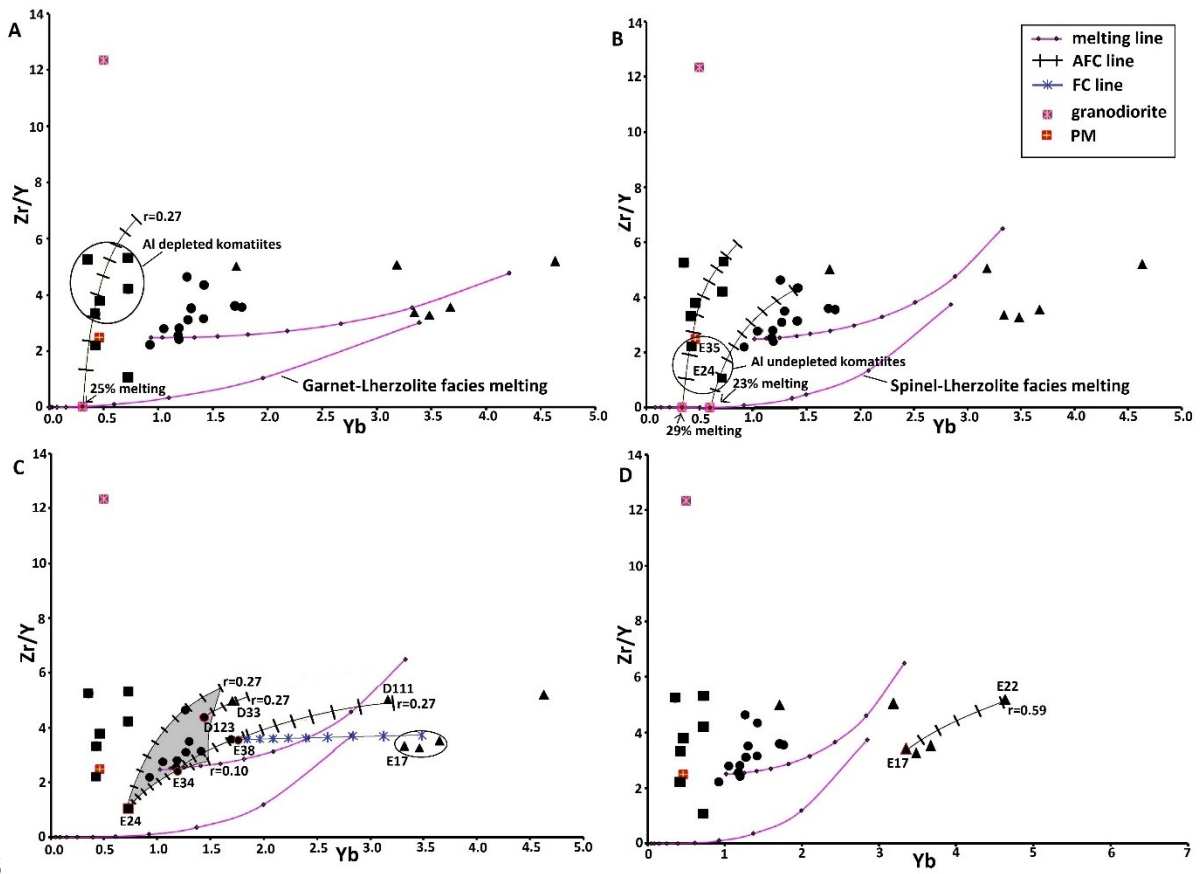


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819**Fig.8.** Bi-variant plot ( $\text{La}/\text{Sm}$  vs.  $\text{La}/\text{Ta}^*$ , after Lassiter and DePaolo, 1997) of the BGB  
820komatiites and basalts to discriminate processes of mixing with continental lithospheric  
821mantle (CLM) and direct contamination with continental crust. Grey arrow shows the overall  
822trend of komatiites and basalts.  $\text{Ta}^* = \text{Nb}/17$  ( $\text{Nb}/\text{Ta} = 17$ , based on a primitive mantle value by  
823Sun and McDonough, 1989). UC~ upper continental crust, LC~ lower continental crust.



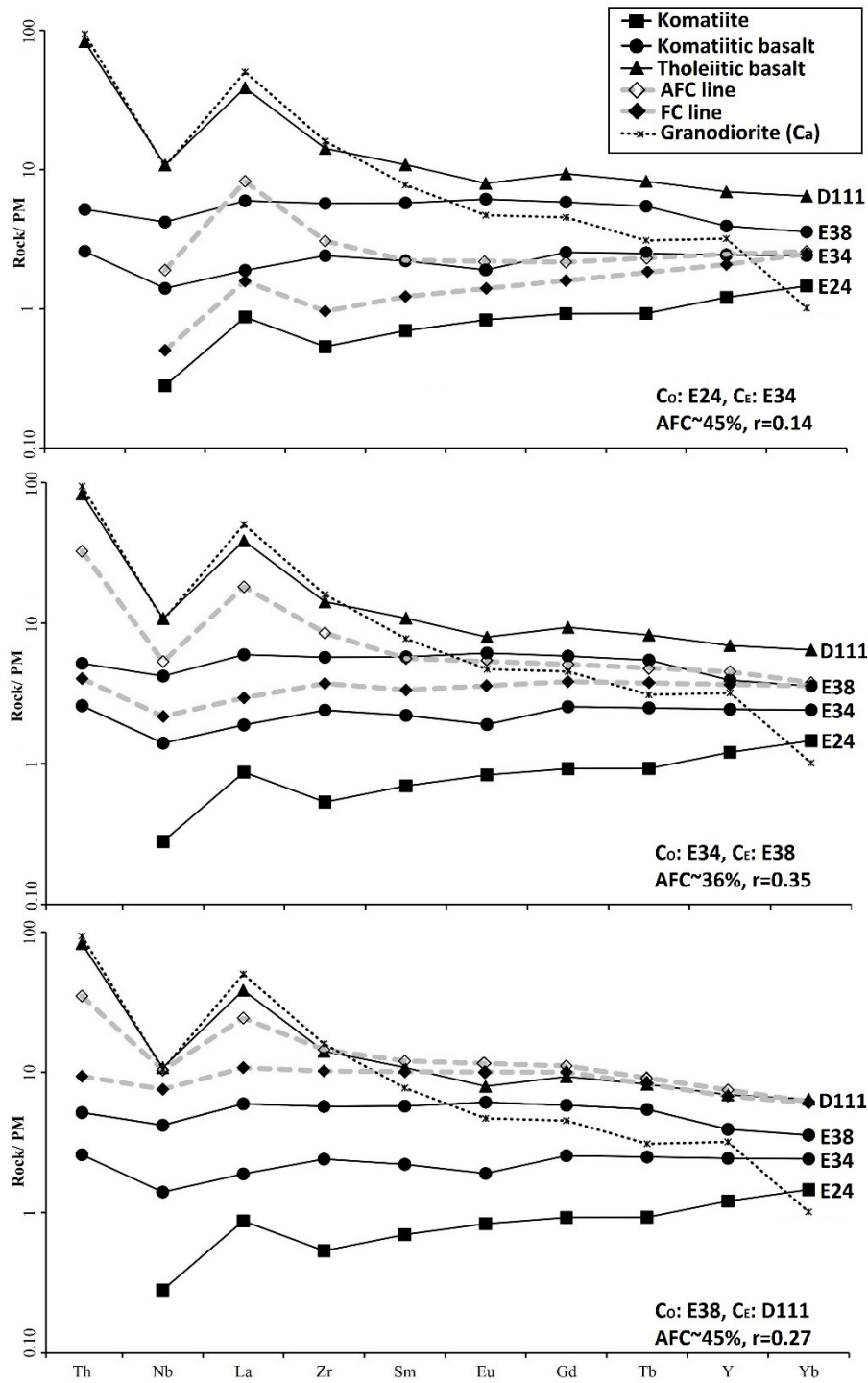
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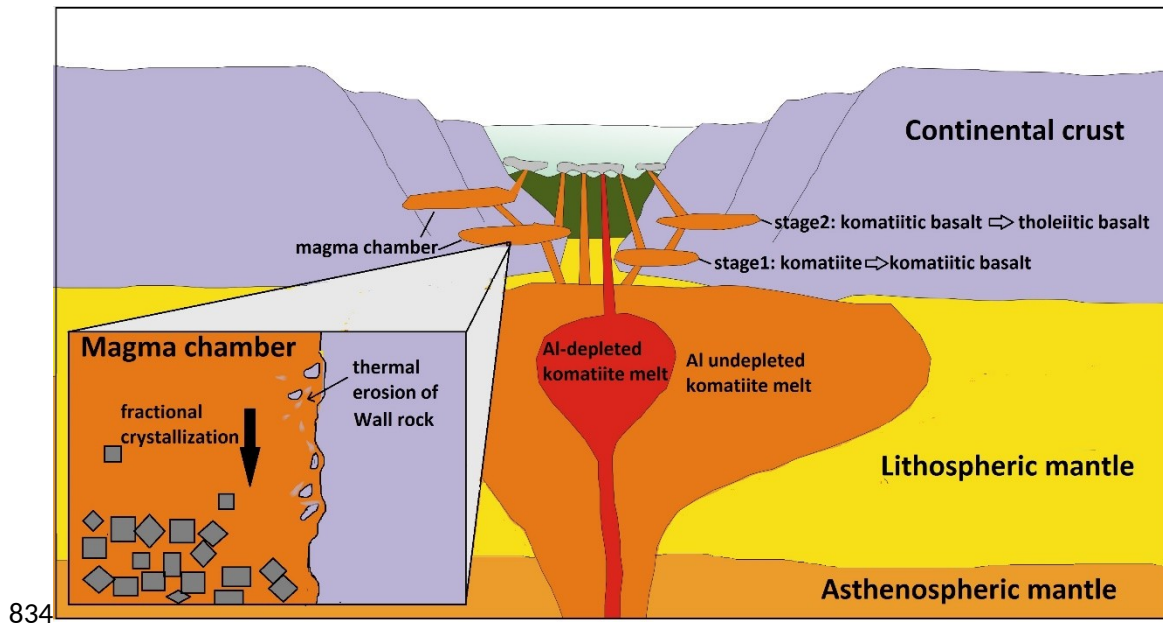
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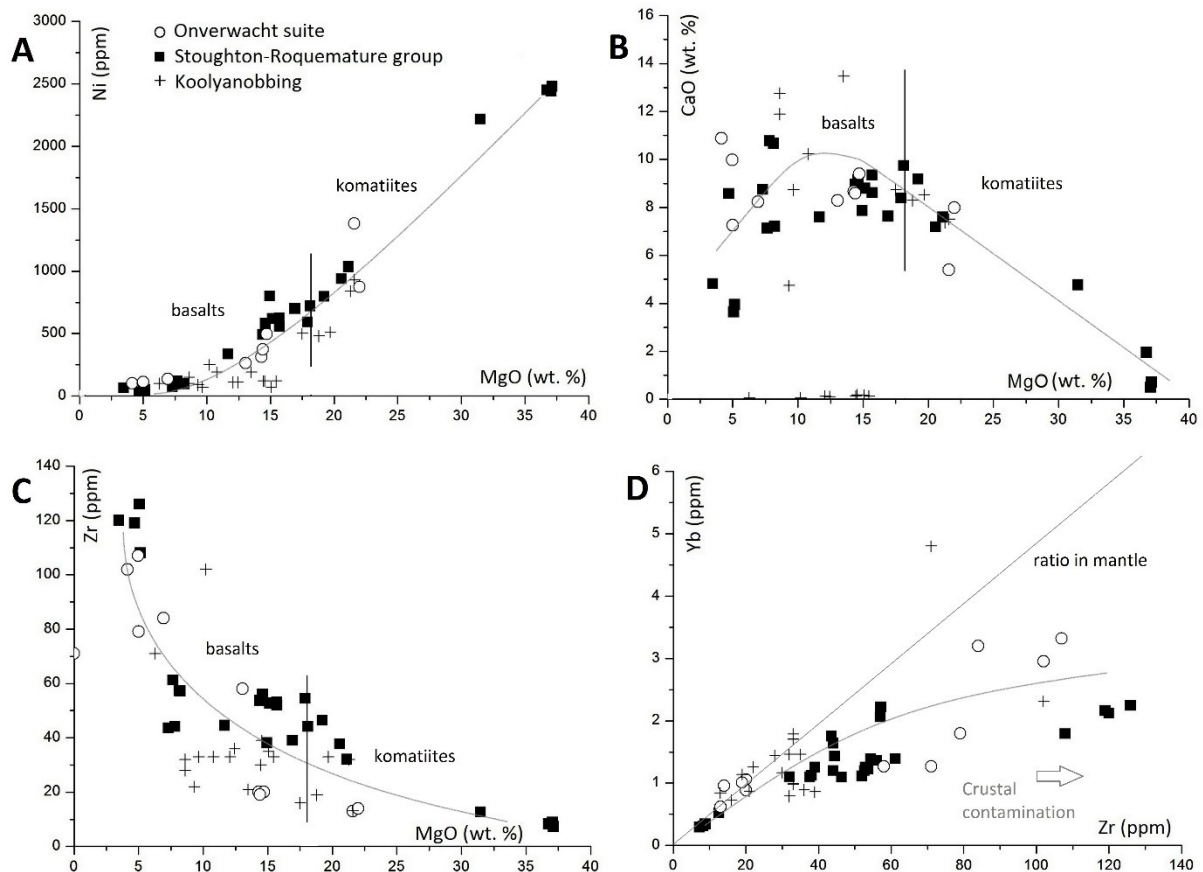
826**Fig. 9.** Melting, assimilation and fractional crystallization modelling of the BGB komatiites  
 827and basalts. The rock type symbols are same as in Fig. 8.

828



**Fig. 10.** Assimilation and fractional crystallization (AFC) modelling of selected BGB metavolcanic rocks.  $C_0$  ~ initial melt composition,  $C_E$  ~ evolved melt composition,  $C_a$  ~ composition of the contaminant, PM ~ Primitive Mantle (Palme and O'Neill, 2014).





838

839 **Fig. 12.** Bi-variant plots (MgO vs. Ni, CaO and Zr, and Zr vs. Yb) of the komatiites and  
 840 basalts present in Stoughton-Roquemature group (~2.7 Ga), Abitibi Greenstone belt,  
 841 Canada (Dostal and Mueller, 2013), Koolyanobbing greenstone belt (~3.0 Ga), Yilgarn  
 842 Craton, Australia (Angerer et al., 2013), and Hooggenoeg Complex, Onverwacht suite  
 843 (~3.5-3.3 Ga), Barberton greenstone belt, South Africa (Furnes et al., 2012). The plots  
 844 show continuous data array of the komatiites and basalts indicated by the projection of  
 845 best fit polynomial trends. Grey straight line in Fig. 11D marks the elemental ratios in  
 846 Earth's mantle (Plame and O'Neill, 2014).