Search for Heavy Long-Lived Particles that Decay to Photons at CDF II

A. Abulencia,²⁴ J. Adelman,¹³ T. Affolder,¹⁰ T. Akimoto,⁵⁵ M. G. Albrow,¹⁷ S. Amerio,⁴³ D. Amidei,³⁵ A. Anastassov,⁵² K. Anikeev,¹⁷ A. Annovi,¹⁹ J. Antos,¹⁴ M. Aoki,⁵⁵ G. Apollinari,¹⁷ T. Arisawa,⁵⁷ A. Artikov,¹⁵ W. Ashmanskas,¹⁷ K. Ankeev, A. Annovi, J. Antos, M. Aoki, G. Apoinnari, T. Arisawa, A. Arukov, W. Asinnańskas,
A. Attal, ³ A. Aurisano, ⁴² F. Azfar, ⁴² P. Azzi-Bacchetta, ⁴³ P. Azzurri, ⁴⁶ N. Bacchetta, ⁴³ W. Badgett, ¹⁷ A. Barbaro-Galtieri, ²⁹
V. E. Barnes, ⁴⁸ B. A. Barnett, ²⁵ S. Baroiant, ⁷ V. Bartsch, ³¹ G. Bauer, ³³ P.-H. Beauchemin, ³⁴ F. Bedeschi, ⁴⁶ S. Behari, ²⁵ G. Bellettini, ⁴⁶ J. Bellinger, ⁵⁹ A. Belloni, ³³ D. Benjamin, ¹⁶ A. Beretvas, ¹⁷ J. Beringer, ²⁹ T. Berry, ³⁰ A. Bhatti, ⁵⁰
M. Binkley, ¹⁷ D. Bisello, ⁴³ I. Bizjak, ³¹ R. E. Blair, ² C. Blocker, ⁶ B. Blumenfeld, ²⁵ A. Bocci, ¹⁶ A. Bodek, ⁴⁹ V. Boisvert, ⁴⁹ G. Bolla, ⁴⁸ A. Bolshov, ³³ D. Bortoletto, ⁴⁸ J. Boudreau, ⁴⁷ A. Boveia, ¹⁰ B. Brau, ¹⁰ L. Brigliadori, ⁵ C. Bromberg, ³⁶ E. Brubaker,¹³ J. Budagov,¹⁵ H. S. Budd,⁴⁹ S. Budd,²⁴ K. Burkett,¹⁷ G. Busetto,⁴³ P. Bussey,²¹ A. Buzatu,³⁴ K. L. Byrum,² S. Cabrera,^{16,q} M. Campanelli,²⁰ M. Campbell,³⁵ F. Canelli,¹⁷ A. Canepa,⁴⁵ S. Carillo,^{18,i} D. Carlsmith,⁵⁹ R. Carosi,⁴⁶ S. Carron,³⁴ B. Casal,¹¹ M. Casarsa,⁵⁴ A. Castro,⁵ P. Catastini,⁴⁶ D. Cauz,⁵⁴ M. Cavalli-Sforza,³ A. Cerri,²⁹ L. Cerrito,^{31,m} S. Carlon, B. Casal, M. Casarsa, A. Castro, F. Catastini, D. Cadz, M. Cavani-Stolza, A. Cern, E. Cernto,
S. H. Chang,²⁸ Y. C. Chen, ¹ M. Chertok, ⁷ G. Chiarelli,⁴⁶ G. Chlachidze,¹⁷ F. Chlebana,¹⁷ I. Cho,²⁸ K. Cho,²⁸ D. Chokheli,¹⁵ J. P. Chou,²² G. Choudalakis,³³ S. H. Chuang,⁵² K. Chung,¹² W. H. Chung,⁵⁹ Y. S. Chung,⁴⁹ M. Cilijak,⁴⁶ C. I. Ciobanu,²⁴ M. A. Ciocci,⁴⁶ A. Clark,²⁰ D. Clark,⁶ M. Coca,¹⁶ G. Compostella,⁴³ M. E. Convery,⁵⁰ J. Conway,⁷ B. Cooper,³¹ K. Copic,³⁵ M. Cordelli,¹⁹ G. Cortiana,⁴³ F. Crescioli,⁴⁶ C. Cuenca Almenar,^{7,q} J. Cuevas,^{11,1} R. Culbertson,¹⁷ J. C. Cully,³⁵ S. DaRonco,⁴³ M. Datta,¹⁷ S. D'Auria,²¹ T. Davies,²¹ D. Dagenhart,¹⁷ P. de Barbaro,⁴⁹ S. De Cecco,⁵¹ A. Deisher,²⁹ G. D. Leuchard, A. Deisher,²⁹ S. De Viene,⁴⁶ F. D. W. D. W. P. W. A. Chart,⁴⁶ K. De Viene,⁴⁶ K. De Viene,⁴⁷ K. Chart,⁴⁶ K. De Viene,⁴⁶ K. De Viene,⁴⁷ K. De Viene,⁴⁶ K. De G. De Lentdecker,^{49,c} G. De Lorenzo,³ M. Dell'Orso,⁴⁶ F. Delli Paoli,⁴³ L. Demortier,⁵⁰ J. Deng,¹⁶ M. Deninno,⁵ D. De Pedis, ⁵¹ P. F. Derwent, ¹⁷ G. P. Di Giovanni, ⁴⁴ C. Dionisi, ⁵¹ B. Di Ruzza, ⁵⁴ J. R. Dittmann, ⁴ M. D'Onofrio, ³ C. Dörr, ²⁶ D. De Fedis, F. F. Derwent, C. D. Grovanni, C. Dronst, D. Dr Ruzza, J. R. Drumann, M. D. Onorrio, C. Dorr,
S. Donati, ⁴⁶ P. Dong, ⁸ J. Donini, ⁴³ T. Dorigo, ⁴³ S. Dube, ⁵² J. Efron, ³⁹ R. Erbacher, ⁷ D. Errede, ²⁴ S. Errede, ²⁴ R. Eusebi, ¹⁷
H. C. Fang, ²⁹ S. Farrington, ³⁰ I. Fedorko, ⁴⁶ W. T. Fedorko, ¹³ R. G. Feild, ⁶⁰ M. Feindt, ²⁶ J. P. Fernandez, ³² R. Field, ¹⁸
G. Flanagan, ⁴⁸ R. Forrest, ⁷ S. Forrester, ⁷ M. Franklin, ²² J. C. Freeman, ²⁹ I. Furic, ¹³ M. Gallinaro, ⁵⁰ J. Galyardt, ¹²
J. E. Garcia, ⁴⁶ F. Garberson, ¹⁰ A. F. Garfinkel, ⁴⁸ C. Gay, ⁶⁰ H. Gerberich, ²⁴ D. Gerdes, ³⁵ S. Giagu, ⁵¹ P. Giannetti, ⁴⁶ K. Gibson,⁴⁷ J. L. Gimmell,⁴⁹ C. Ginsburg,¹⁷ N. Giokaris,^{15,a} M. Giordani,⁵⁴ P. Giromini,¹⁹ M. Giunta,⁴⁶ G. Giurgiu,²⁵ V. Glagolev,¹⁵ D. Glenzinski,¹⁷ M. Gold,³⁷ N. Goldschmidt,¹⁸ J. Goldstein,^{42,b} A. Golossanov,¹⁷ G. Gomez,¹¹ K. Gibson, J. L. Gimmell, "C. Ginsburg, "N. Giokaris, ""M. Giordan," 'P. Giromini, "M. Giunta, "G. Giurgu, "V. Glagolev,¹⁵ D. Glenzinski,¹⁷ M. Gold,³⁷ N. Goldschmidt,¹⁸ J. Goldstein,^{42,b} A. Golossanov,¹⁷ G. Gomez,¹¹ G. Gomez,²² C. Grosso-Pilcher,¹³ R. C. Group,¹⁷ U. Grundler,²⁴ J. Guimaraes da Costa,²² Z. Gunay-Unalan,³⁶ C. Haber,²⁹ K. Hahn,³³ S. R. Hahn,¹⁷ E. Halkiadakis,⁵² A. Hamilton,²⁰ B.-Y. Han,⁴⁹ J. Y. Han,⁴⁹ R. Handler,⁵⁹
F. Happacher,¹⁹ K. Hara,⁵⁵ D. Hare,⁵² M. Hare,⁵⁶ S. Harper,⁴² R. F. Harr,⁵⁸ R. M. Harris,¹⁷ M. Hartz,⁴⁷ K. Hatakeyama,⁵⁰ J. Hauser,⁸ C. Hays,⁴² M. Heck,²⁶ A. Heijboer,⁴⁵ B. Heinemann,²⁹ J. Heinrich,⁴⁵ C. Henderson,³³ M. Herndon,⁵⁹ J. Heuser,²⁶ D. Hidas,¹⁶ C. S. Hill,^{10,b} D. Hirschbuehl,²⁶ A. Hocker,¹⁷ A. Holloway,²² S. Hou,¹ M. Houlden,³⁰ S.-C. Hsu,⁹ B. T. Huffman,⁴² R. E. Hughes,³⁹ U. Husemann,⁶⁰ J. Huston,³⁶ J. Incandela,¹⁰ G. Introzzi,⁴⁶ M. Iori,⁵¹ A. Ivanov,⁷ B. Iyutin,³³ E. James,¹⁷ D. Jang,⁵² B. Jayatilaka,¹⁶ D. Jeans,⁵¹ E. J. Jeon,²⁸ S. Jindariani,¹⁸ W. Johnson,⁷ M. Jones,⁴⁸ K. K. Joo,²⁸ S. Y. Jun,¹² J. E. Jung,²⁸ T. R. Junk,²⁴ T. Kamon,⁵³ P.E. Karchin,⁵⁸ Y. Kato,⁴¹ Y. Kemp,²⁶ R. Kephart,¹⁷ U. Kerzel,²⁶ V. Khotilovich,⁵³ B. Kiliminster,³⁹ D. H. Kim,²⁸ H. S. Kim,²⁸ M. J. Kim,¹⁷ S. B. Kim,⁸⁵ S. H. Kim,⁵⁵ Y. K. Kim,¹³ N. Kimura,⁵⁵ L. Kirsch,⁶ S. Klimenko,¹⁸ M. Klute,³³ B. Knuteson,³³ B. R. Ko,¹⁶ K. Kondo,⁵⁷ D. J. Kong,²⁸ J. Konigsberg,¹⁸ A. Korytov,¹⁸ A. V. Kotwal,¹⁶ A. C. Kraan,⁴⁵ J. Kraus,²⁴ M. Kreps,²⁶ J. Kroll,⁴⁵ N. Krunmack,⁴ M. Kruse,¹⁶ V. Krutelyov,¹⁰ T. Kubo,⁵⁵ S. E. Kuhlmann,² T. Kuhr,²⁶ N. P. Kulkarni,⁵⁸ Y. Kusakabe,⁵⁷ S. Kwang,¹³ A. T. Laasanen,⁴⁸ S. Laui,³⁴ S. Lami,⁴⁶ S. Lammel,¹⁷ M. Lancaster,³¹ R. L. Lander,⁷ K. Lannon,³⁹ A. Lath,⁵² G. Latino,⁴⁶ I. Lazzizzera,⁴³ T. LeCompte,² E. Lee,⁵³ J. Lee,⁴⁰ J. Lee,²⁸ S. W. Lee, T. Maruyama,⁵⁵ P. Mastrandrea,⁵¹ T. Masubuchi,⁵⁵ H. Matsunaga,⁵⁵ M. E. Mattson,⁵⁸ R. Mazini,³⁴ P. Mazzanti,⁵ K. S. McFarland,⁴⁹ P. McIntyre,⁵³ R. McNulty,^{30,f} A. Mehta,³⁰ P. Mehtala,²³ S. Menzemer,^{11,i} A. Menzione,⁴⁶ P. Merkel,⁴⁸ C. Mesropian,⁵⁰ A. Messina,³⁶ T. Miao,¹⁷ N. Miladinovic,⁶ J. Miles,³³ R. Miller,³⁶ C. Mills,¹⁰ M. Milnik,²⁶ A. Mitra,¹ G. Mitselmakher,¹⁸ A. Miyamoto,²⁷ S. Moed,²⁰ N. Moggi,⁵ B. Mohr,⁸ C. S. Moon,²⁸ R. Moore,¹⁷ M. Morello,⁴⁶ P. Movilla Fernandez,²⁹ J. Mülmenstädt,²⁹ A. Mukherjee,¹⁷ Th. Muller,²⁶ R. Mumford,²⁵ P. Murat,¹⁷ M. Mussini,⁵

J. Nachtman,¹⁷ A. Nagano,⁵⁵ J. Naganoma,⁵⁷ K. Nakamura,⁵⁵ I. Nakano,⁴⁰ A. Napier,⁵⁶ V. Necula,¹⁶ C. Neu,⁴⁵ M. S. Neubauer,⁹ J. Nielsen,^{29,n} L. Nodulman,² O. Norniella,³ E. Nurse,³¹ S. H. Oh,¹⁶ Y. D. Oh,²⁸ I. Oksuzian,¹⁸ T. Okusawa,⁴¹ R. Oldeman,³⁰ R. Orava,²³ K. Osterberg,²³ C. Pagliarone,⁴⁶ E. Palencia,¹¹ V. Papadimitriou,¹⁷ T. Okusawa,⁴¹ R. Oldeman,³⁰ R. Orava,²³ K. Osterberg,²³ C. Pagliarone,⁴⁶ E. Palencia,¹¹ V. Papadimitriou,¹⁷
A. Papaikonomou,²⁶ A. A. Paramonov,¹³ B. Parks,³⁹ S. Pashapour,³⁴ J. Patrick,¹⁷ G. Pauletta,⁵⁴ M. Paulini,¹² C. Paus,³³
D. E. Pellett,⁷ A. Penzo,⁵⁴ T. J. Phillips,¹⁶ G. Piacentino,⁴⁶ J. Piedra,⁴⁴ L. Pinera,¹⁸ K. Pitts,²⁴ C. Plager,⁸ L. Pondrom,⁵⁹
X. Portell,³ O. Poukhov,¹⁵ N. Pounder,⁴² F. Prakoshyn,¹⁵ A. Pronko,¹⁷ J. Proudfoot,² F. Ptohos,^{19,e} G. Punzi,⁴⁶ J. Pursley,²⁵
J. Rademacker,^{42,b} A. Rahaman,⁴⁷ V. Ramakrishnan,⁵⁹ N. Ranjan,⁴⁸ I. Redondo,³² B. Reisert,¹⁷ V. Rekovic,³⁷ P. Renton,⁴²
M. Rescigno,⁵¹ S. Richter,²⁶ F. Rimondi,⁵ L. Ristori,⁴⁶ A. Robson,²¹ T. Rodrigo,¹¹ E. Rogers,²⁴ S. Rolli,⁵⁶ R. Roser,¹⁷
M. Rossi,⁵⁴ R. Rossin,¹⁰ P. Roy,³⁴ A. Ruiz,¹¹ J. Russ,¹² V. Rusu,¹³ H. Saarikko,²³ A. Safonov,⁵³ W. K. Sakumoto,⁴⁹
G. Salamanna,⁵¹ O. Saltó,³ L. Santi,⁵⁴ S. Sarkar,⁵¹ L. Sartori,⁴⁶ K. Sato,¹⁷ P. Savard,³⁴ A. Savoy-Navaro,⁴⁴ T. Scheidle,²⁶
P. Schlabach,¹⁷ E. E. Schmidt,¹⁷ M. P. Schmidt,⁶⁰ M. Schmitt,³⁸ T. Schwarz,⁷ L. Scodellaro,¹¹ A. L. Scott,¹⁰ A. Scribano,⁴⁶
F. Scuri,⁴⁶ A. Sedov,⁴⁸ S. Seidel,³⁷ Y. Seiya,⁴¹ A. Semenov,¹⁵ L. Sexton-Kennedy,¹⁷ A. Sfyrla,²⁰ S. Z. Shalhout,⁵⁸
M. D. Shapiro,²⁹ T. Shears,³⁰ P. F. Shepard,⁴⁷ D. Sherman,²² M. Shimojima,^{55,k} M. Shochet,¹³ Y. Shop,⁵⁹ I. Shrevber,²⁰ M. D. Shapiro,²⁹ T. Shears,³⁰ P. F. Shepard,⁴⁷ D. Sherman,²² M. Shimojima,^{55,k} M. Shochet,¹³ Y. Shon,⁵⁹ I. Shreyber,²⁰ A. Sidoti,⁴⁶ P. Sinervo,³⁴ A. Sisakyan,¹⁵ A. J. Slaughter,¹⁷ J. Slaunwhite,³⁹ K. Sliwa,⁵⁶ J. R. Smith,⁷ F. D. Snider,¹⁷ R. Snihur,³⁴ M. Soderberg,³⁵ A. Soha,⁷ S. Somalwar,⁵² V. Sorin,³⁶ J. Spalding,¹⁷ F. Spinella,⁴⁶ T. Spreitzer,³⁴ P. Squillacioti,⁴⁶ M. Stanitzki,⁶⁰ A. Staveris-Polykalas,⁴⁶ R. St. Denis,²¹ B. Stelzer,⁸ O. Stelzer-Chilton,⁴² D. Stentz,³⁸ P. Squillacioti,⁴⁶ M. Stanitzki,⁶⁰ A. Staveris-Polykalas,⁴⁶ R. St. Denis,²¹ B. Stelzer,⁸ O. Stelzer-Chilton,⁴² D. Stentz,³⁸ J. Strologas,³⁷ D. Stuart,¹⁰ J. S. Suh,²⁸ A. Sukhanov,¹⁸ H. Sun,⁵⁶ I. Suslov,¹⁵ T. Suzuki,⁵⁵ A. Taffard,^{24,p} R. Takashima,⁴⁰ Y. Takeuchi,⁵⁵ R. Tanaka,⁴⁰ M. Tecchio,³⁵ P.K. Teng,¹ K. Terashi,⁵⁰ J. Thom,^{17,d} A. S. Thompson,²¹ E. Thomson,⁴⁵ P. Tipton,⁶⁰ V. Tiwari,¹² S. Tkaczyk,¹⁷ D. Toback,⁵³ S. Tokar,¹⁴ K. Tollefson,³⁶ T. Tomura,⁵⁵ D. Tonelli,⁴⁶ S. Torre,¹⁹ D. Torretta,¹⁷ S. Tourneur,⁴⁴ W. Trischuk,³⁴ S. Tsuno,⁴⁰ Y. Tu,⁴⁵ N. Turini,⁴⁶ F. Ukegawa,⁵⁵ S. Uozumi,⁵⁵ S. Vallecorsa,²⁰ N. van Remortel,²³ A. Varganov,³⁵ E. Vataga,³⁷ F. Vazquez,^{18,i} G. Velev,¹⁷ G. Veramendi,²⁴ V. Veszpremi,⁴⁸ M. Vidal,³² R. Vidal,¹⁷ I. Vila,¹¹ R. Vilar,¹¹ T. Vine,³¹ I. Vollrath,³⁴ I. Volobouev,^{29,o} G. Volpi,⁴⁶ F. Würthwein,⁹ P. Wagner,⁵³ R. G. Wagner,² R. L. Wagner,¹⁷ J. Wagner,²⁶ W. Wagner,²⁶ R. Wallny,⁸ S. M. Wang,¹ A. Warburton,³⁴ D. Waters,³¹ M. Weinberger,⁵³ W. C. Wester III,¹⁷ B. Whitehouse,⁵⁶ D. Whiteson,⁴⁵ A. B. Wicklund,² E. Wicklund,¹⁷ G. Williams,³⁴ H. H. Williams,⁴⁵ P. Wilson,¹⁷ B. L. Winer,³⁹ P. Wittich,^{17,d} S. Wolbers,¹⁷ C. Wolfe,¹³ T. Wright,³⁵ X. Wu,²⁰ S. M. Wynne,³⁰ A. Yagil,⁹ K. Yamamoto,⁴¹ J. Yamaoka,⁵² T. Yamashita,⁴⁰ C. Yang,⁶⁰ U. K. Yang,^{13,j} Y.C. Yang,²⁸ W. M. Yao,²⁹ G. P. Yeh,¹⁷ J. Yoh,¹⁷ K. Yorita,¹³ T. Yoshida,⁴¹ G. B. Yu,⁴⁹ I. Yu,²⁸ S.S. Yu,¹⁷ J. C. Yun,¹⁷ L. Zanello,⁵¹ A. Zanetti,⁵⁴ I. Zaw,²² X. Zhang,²⁴ J. Zhou,⁵² and S. Zucchelli⁵

(CDF Collaboration)^a

¹Institute of Physics, Academia Sinica, Taipei, Taiwan 11529, Republic of China

²Argonne National Laboratory, Argonne, Illinois 60439, USA

³Institut de Fisica d'Altes Energies, Universitat Autonoma de Barcelona, E-08193, Bellaterra (Barcelona), Spain

⁴Baylor University, Waco, Texas 76798, USA

⁵Istituto Nazionale di Fisica Nucleare, University of Bologna, I-40127 Bologna, Italy

⁶Brandeis University, Waltham, Massachusetts 02254, USA

⁷University of California, Davis, Davis, California 95616, USA

⁸University of California, Los Angeles, Los Angeles, California 90024, USA

⁹University of California, San Diego, La Jolla, California 92093, USA

¹⁰University of California, Santa Barbara, Santa Barbara, California 93106, USA

¹¹Instituto de Fisica de Cantabria, CSIC-University of Cantabria, 39005 Santander, Spain

¹²Carnegie Mellon University, Pittsburgh, Pennsylvania 15213, USA

¹³Enrico Fermi Institute, University of Chicago, Chicago, Illinois 60637, USA

¹⁴Comenius University, 842 48 Bratislava, Slovakia;

Institute of Experimental Physics, 040 01 Kosice, Slovakia

¹⁵Joint Institute for Nuclear Research, RU-141980 Dubna, Russia

¹⁶Duke University, Durham, North Carolina 27708, USA

¹⁷Fermi National Accelerator Laboratory, Batavia, Illinois 60510, USA

¹⁸University of Florida, Gainesville, Florida 32611, USA

¹⁹Laboratori Nazionali di Frascati, Istituto Nazionale di Fisica Nucleare, I-00044 Frascati, Italy

²⁰University of Geneva, CH-1211 Geneva 4, Switzerland

²¹Glasgow University, Glasgow G12 8QQ, United Kingdom

²²Harvard University, Cambridge, Massachusetts 02138, USA

PRL 99, 121801 (2007) 21 SEPTEMBER 2007 ²³Division of High Energy Physics, Department of Physics, University of Helsinki and Helsinki Institute of Physics, FIN-00014, Helsinki, Finland ²⁴University of Illinois, Urbana, Illinois 61801, USA ²⁵The Johns Hopkins University, Baltimore, Maryland 21218, USA ²⁶Institut für Experimentelle Kernphysik, Universität Karlsruhe, 76128 Karlsruhe, Germany ²⁷High Energy Accelerator Research Organization (KEK), Tsukuba, Ibaraki 305, Japan ²⁸Center for High Energy Physics: Kyungpook National University, Taegu 702-701, Korea; Seoul National University, Seoul 151-742, Korea; SungKyunKwan University, Suwon 440-746, Korea ²⁹Ernest Orlando Lawrence Berkeley National Laboratory, Berkeley, California 94720, USA ³⁰University of Liverpool, Liverpool L69 7ZE, United Kingdom ³¹University College London, London WC1E 6BT, United Kingdom ³²Centro de Investigaciones Energeticas Medioambientales y Tecnologicas, E-28040 Madrid, Spain ³Massachusetts Institute of Technology, Cambridge, Massachusetts 02139, USA ³⁴Institute of Particle Physics: McGill University, Montréal, Canada H3A 2T8; and University of Toronto, Toronto, Canada M5S 1A7 ³⁵University of Michigan, Ann Arbor, Michigan 48109, USA ³⁶Michigan State University, East Lansing, Michigan 48824, USA ³⁷University of New Mexico, Albuquerque, New Mexico 87131, USA ³⁸Northwestern University, Evanston, Illinois 60208, USA ³⁹The Ohio State University, Columbus, Ohio 43210, USA ⁴⁰Okayama University, Okayama 700-8530, Japan ⁴¹Osaka City University, Osaka 588, Japan ⁴²University of Oxford, Oxford OX1 3RH, United Kingdom ⁴³University of Padova, Istituto Nazionale di Fisica Nucleare, Sezione di Padova-Trento, I-35131 Padova, Italy ¹⁴LPNHE, Universite Pierre et Marie Curie/IN2P3-CNRS, UMR7585, Paris, F-75252 France ⁴⁵University of Pennsylvania, Philadelphia, Pennsylvania 19104, USA ⁴⁶Istituto Nazionale di Fisica Nucleare Pisa, Universities of Pisa, Siena and Scuola Normale Superiore, I-56127 Pisa, Italy ⁴⁷University of Pittsburgh, Pittsburgh, Pennsylvania 15260, USA ⁴⁸Purdue University, West Lafavette, Indiana 47907, USA ⁴⁹University of Rochester, Rochester, New York 14627, USA ⁵⁰The Rockefeller University, New York, New York 10021, USA ⁵¹Istituto Nazionale di Fisica Nucleare, Sezione di Roma 1, University of Rome "La Sapienza," I-00185 Roma, Italy ⁵²Rutgers University, Piscataway, New Jersey 08855, USA ⁵³Texas A&M University, College Station, Texas 77843, USA ⁵⁴Istituto Nazionale di Fisica Nucleare, University of Trieste/Udine, Italy ⁵⁵University of Tsukuba, Tsukuba, Ibaraki 305, Japan ⁵⁶Tufts University, Medford, Massachusetts 02155, USA ⁵⁷Waseda University, Tokyo 169, Japan ⁵⁸Wayne State University, Detroit, Michigan 48201, USA ⁵⁹University of Wisconsin, Madison, Wisconsin 53706, USA ⁶⁰Yale University, New Haven, Connecticut 06520, USA (Received 5 April 2007; published 17 September 2007) We present the first search for heavy, long-lived particles that decay to photons at a hadron collider. We

use a sample of γ + jet + missing transverse energy events in $p\bar{p}$ collisions at $\sqrt{s} = 1.96$ TeV taken with the CDF II detector. Candidate events are selected based on the arrival time of the photon at the detector. Using an integrated luminosity of 570 pb^{-1} of collision data, we observe 2 events, consistent with the background estimate of 1.3 ± 0.7 events. While our search strategy does not rely on model-specific dynamics, we set cross section limits in a supersymmetric model with $\widetilde{\chi}_1^0 \to \gamma \widetilde{G}$ and place the world-best 95% C.L. lower limit on the $\widetilde{\chi}_1^0$ mass of 101 GeV/ c^2 at $\tau_{\widetilde{\chi}_1^0} = 5$ ns.

DOI: 10.1103/PhysRevLett.99.121801

PACS numbers: 13.85.Rm, 12.60.Jv, 13.85.Qk, 14.80.Ly

Searches for events with final state photons and missing transverse energy $(\not\!\!\!E_T)$ at collider experiments are sensitive to new physics from a wide variety of models [1] including gauge mediated supersymmetry breaking (GMSB) [2]. In these models the lightest neutralino $(\widetilde{\chi}_1^0)$ decays into a

photon (γ) and a weakly interacting, stable gravitino (G)that gives rise to $\not\!\!\!E_T$ by leaving the detector without depositing any energy. The observation of an $ee\gamma\gamma \not\!\!\!/ E_T$ candidate event by the CDF experiment during run I at the Fermilab Tevatron [3] has increased the interest in experimental tests of this class of theories. Most subsequent searches have focused on promptly produced photons [4,5]; however, the $\tilde{\chi}_1^0$ can have a lifetime on the order of nanoseconds or more. This Letter summarizes [6] the first search for heavy, long-lived particles that decay to photons at a hadron collider. The data comprise 570 ± 34 pb⁻¹ of $p\bar{p}$ collisions at $\sqrt{s} = 1.96$ TeV from the Tevatron collected with the CDF II detector [7]. Previous searches for nanosecond-lifetime particles using nontiming techniques have yielded null results [8].

We optimize our selection requirements using a GMSB model with a standard choice of parameters [9] and vary the values of the $\tilde{\chi}_1^0$ mass and lifetime. However, the final search strategy is chosen to be sufficiently general and independent of the specific GMSB model dynamics to yield results that are approximately valid for any model producing the same reconstructed final state topology and kinematics [10]. The inclusive GMSB production cross section is dominated by pair production of gauginos. The gauginos decay promptly, resulting in a pair of long-lived $\widetilde{\chi}_1^0$'s in association with other final state particles that can be identified as jets. For a heavy $\widetilde{\chi}_1^0$ decaying inside the detector, the photon can arrive at the face of the detector with a time delay relative to promptly produced photons. To have good sensitivity for nanosecond-lifetime $\tilde{\chi}_1^0$'s [10], we search for events that contain a time-delayed photon, $\not\!\!\!E_T$, and ≥ 1 jet.

Here we briefly describe the aspects of the detector relevant to this analysis. The magnetic spectrometer consists of tracking devices inside a 3-m diameter, 5-m long superconducting solenoid magnet that operates at 1.4 T. An eight-layer silicon microstrip detector array and a 3.1-m long drift chamber with 96 layers of sense wires measure the position (\vec{x}_i) and time (t_i) of the $p\bar{p}$ interaction [11] and the momenta of charged particles. Muons from collisions or cosmic rays are identified by a system of drift chambers situated outside the calorimeters in the region with pseudorapidity $|\eta| < 1.1$. The calorimeter consists of projective towers with electromagnetic and hadronic compartments. It is divided into a central barrel that surrounds the solenoid coil ($|\eta| < 1.1$) and a pair of end plugs that cover the region $1.1 < |\eta| < 3.6$. Both calorimeters are used to identify and measure the energy and position of photons, electrons, jets, and $\not\!\!\!E_T$. The electromagnetic calorimeters were recently instrumented with a new system, the EMTiming system (completed in Fall 2004) [12], that measures the arrival time of electrons and photons in each tower with $|\eta| < 2.1$ for all energies above \sim 5 GeV. The system has a fixed-threshold discriminator and a time-to-digital converter (TDC) connected to each electromagnetic tower and has an intrinsic time-of-arrival resolution of 0.6 ns.

The time and position of arrival of the photon at the calorimeter, t_f and \vec{x}_f , are used to separate the photons from the decays of heavy, long-lived $\tilde{\chi}_1^0$'s from promptly

produced photons or photons from noncollision sources. We define the corrected arrival time of the photon as $t_c^{\gamma} \equiv t_f - t_i - \frac{|\vec{x}_f - \vec{x}_i|}{c}$. The t_c^{γ} distribution for promptly produced, high energy photons is Gaussian with a mean of zero by construction and has a standard deviation that depends only on the measurement resolution assuming that the $p\bar{p}$ production vertex has been correctly identified. Photons from heavy, long-lived particles can have arrival times that are many standard deviations larger than zero.

The analysis preselection is summarized in Table I, with the effect on an example GMSB model point at $m_{\tilde{\chi}_1^0} =$ 100 GeV/ c^2 and $au_{ ilde{\chi}_1^0} = 5$ ns for comparison. It begins with events passing an online, three-level trigger by having a photon candidate in the region $|\eta| < 1.1$ with $E_T >$ 25 GeV and $\not\!\!E_T > 25$ GeV. Offline, the highest E_T photon candidate in the fiducial region of the calorimeter is required to have $E_T > 30$ GeV and to pass the standard photon identification requirements [4] with a minor modification [13]. We require the event to have $\not\!\!\!E_T > 30 \text{ GeV}$ where the trigger is 100% efficient. We require at least one jet with $|\eta^{\text{jet}}| < 2.0$ and $E_T^{\text{jet}} > 30$ GeV [14]. Since a second photon can be identified as a jet, the analysis is sensitive to signatures where one or both $\tilde{\chi}_1^0$'s decay inside the detector. To ensure a high quality t_i and \vec{x}_i measurement, we require a vertex with at least 4 tracks, $\sum_{\text{tracks}} p_T >$ 15 GeV/c, and $|z_i| < 60$ cm; this also helps to reduce noncollision backgrounds. For events with multiple reconstructed vertices, we pick the vertex with the highest $\sum_{\text{tracks}} p_T$. To reduce cosmic ray background, events are rejected if there are hits in a muon chamber that are not matched to any track and are within 30° of the photon. After the above requirements there are 11 932 events in the data sample.

TABLE I. The data selection criteria and efficiencies for the example GMSB model point. The efficiencies listed are, in general, model-dependent and have a fractional uncertainty of 10%. Model-independent efficiencies are indicated with an asterisk. The collision fiducial requirement of $|z_i| < 60$ cm is part of the good vertex requirement (95%) and is estimated from data.

Requirements	Cumulative (individual) efficiency (%)
Preselection requirements	
$E_T^{\gamma} > 30 \text{ GeV}, \not\!\!\!E_T > 30 \text{ GeV}$	54 (54)
Photon ID and fiducial, $ \eta < 1.0$	39 (74)*
Good vertex, $\sum_{\text{tracks}} p_T > 15 \text{ GeV}/c$	31 (79)
$ \eta^{\rm jet} < 2.0, E_T^{\rm jet} > 30 { m GeV}$	24 (77)
Cosmic ray rejection	23 (98)*
Requirements after optimization	
$\not\!\!\!E_T > 40 \text{ GeV}, E_T^{\text{jet}} > 35 \text{ GeV}$	21 (92)
$\Delta \phi(\not\!\!E_T, \text{ jet}) > 1 \text{ rad}$	18 (86)
$2 \text{ ns} < t_c^{\gamma} < 10 \text{ ns}$	6 (33)

There are two major classes of background events: collision and noncollision photon candidates. Collision photons are presumed to come from standard model interactions, e.g., $\gamma + \text{jet} + \text{mismeasured } \not E_T$, dijet + mismeasured $\not E_T$, where the jet is misidentified as a γ , and $W \rightarrow e\nu$, where the electron is misidentified as a γ . Noncollision backgrounds come from cosmic rays and beam effects that can produce photon candidates, $\not E_T$, and sometimes the reconstructed jet. We separate data events as a function of t_c^{γ} into several control regions that allow us to estimate the number of background events in the final signal region by fitting to the data using collision and noncollision shapes as shown in Fig. 1.

Collision photons are subdivided in two subclasses: correct and incorrect vertex selection [12]. An incorrect vertex can be selected when two or more collisions occur in one beam bunch crossing, making it possible that the highest reconstructed $\sum_{\text{tracks}} p_T$ vertex does not produce the photon. While the fraction of events with incorrect vertices depends on the final event selection criteria, the t_c^{γ} distribution for each subclass is estimated separately using $W \rightarrow e\nu$ data where the electron track is dropped from the vertexing. For events with a correctly associated vertex, the t_c^{γ} distribution is Gaussian and centered at zero with a standard deviation of 0.64 ns [12]. For those with an incorrectly selected vertex the t_c^{γ} distribution is also Gaussian with a standard deviation of 2.05 ns.

The t_c^{γ} distributions for both noncollision backgrounds are estimated separately from data using events with no reconstructed tracks. Photon candidates from cosmic rays are not correlated in time with collisions, and therefore their t_c^{γ} distribution is roughly flat. Beam halo photon candidates are produced by muons that originate upstream of the detector (from the *p* direction) and travel through the calorimeter, typically depositing small amounts of energy. When the muon deposits significant energy in the EM calorimeter, it can be misidentified as a photon and cause



FIG. 1 (color online). The time distribution for photons passing all but the final timing requirement for the background predictions, data (marked with a star for the signal region), and a GMSB signal for the example model point.

The background prediction uses control regions outside the signal time window but well within the 132 ns time window that the calorimeter uses to measure the energy. The noncollision backgrounds are normalized to match the number of events in two time windows: a beam halodominated window at $\{-20, -6\}$ ns, selected to be 3σ away from the wrong vertex collision background, and a cosmic rays-dominated window at {25, 90} ns, well away from the standard model and beam halo contributions. The collision background is estimated by fitting events in the $\{-10, 1.2\}$ ns window with the noncollision contribution subtracted and with the fraction of correct to incorrect vertex events allowed to vary. In this way the background for the signal region is entirely estimated from data samples. The systematic uncertainty on the background estimate is dominated by our ability to calibrate the mean of the t_c^{γ} distribution for prompt photons. We find a variation of 200 ps on the mean and 20 ps on the standard deviation of the distribution by considering various possible event selection criteria. The systematic uncertainty due to these variations is added in quadrature with the statistical uncertainties of the final fit procedure.

We estimate the sensitivity to heavy, long-lived particles that decay to photons using GMSB models for different χ_1^0 masses and lifetimes. Events from all SUSY processes are simulated with the PYTHIA Monte Carlo program [15] along with the detector simulation [16]. The acceptance is the ratio of simulated events that pass all the requirements to all events produced. It is used in the optimization procedure and in the final limit setting and depends on a number of effects. The fraction of $\widetilde{\chi}_1^0$ decays in the detector volume is the dominant effect on the acceptance. For a given lifetime this depends on the boost of the χ_1^0 . A highly boosted $\widetilde{\chi}_1^0$ that decays in the detector typically does not contribute to the acceptance because it tends to produce a photon traveling in the same direction as the $\tilde{\chi}_1^0$. Thus, the photon's arrival time is indistinguishable from promptly produced photons. At small boosts the decay is more likely to happen inside the detector, and the decay angle is more likely to be large, which translates into a larger delay for the photon. The fraction of events with a delayed photon arrival time initially rises as a function of $\tilde{\chi}_1^0$ lifetime, but falls as the fraction of $\tilde{\chi}_1^0$'s decaying outside the detector begins to dominate. In the $\tilde{\chi}_1^0$ mass region consid-ered ($65 \le m_{\tilde{\chi}_1^0} \le 150 \text{ GeV}/c^2$), the acceptance peaks at a lifetime of around 5 ns. The acceptance also depends on the mass as the boost effects are mitigated by the ability to produce high energy photons or $\not\!\!\!E_T$ in the collision [10].

The total systematic uncertainty of 10% on the acceptance is dominated by the uncertainty on the mean of the t_c^{γ} distribution (7%) and on the photon ID efficiency (5%). Other significant contributions come from uncertainties on



FIG. 2. The contours of constant 95% C.L. upper cross section limits for a GMSB model [9].

initial and final state radiation (3%), jet energy measurement (3%), and the parton distribution functions (1%).

We determine the kinematic and t_c^{γ} selection requirements that define the final data sample by optimizing the predicted cross section limit without looking at the data in the signal region. To compute the predicted 95% confidence level (C.L.) cross section upper limit [17], we combine the predicted GMSB signal and background estimates with the systematic uncertainties using a Bayesian method [18]. The predicted limits are optimized by simultaneously varying the selection requirements for $\not\!\!E_T$, photon E_T , jet $(\Delta \phi(\not\!\!E_T, \text{jet}))$, and t_c^{γ} . The $\Delta \phi(\not\!\!E_T, \text{jet})$ requirement rejects events where the $\not\!\!\!E_T$ is overestimated because of a poorly measured jet. While each point in the considered $\tilde{\chi}_1^0$ lifetime vs mass space gives a slightly different optimization, we choose a single set of requirements because it simplifies the final analysis, while only causing a <4% loss of sensitivity. The optimized requirements are summarized in Table I. The acceptance for the example model point is estimated to be $(6.3 \pm 0.6)\%$.

After all kinematic requirements, 508 events are observed in the data before the final signal region time requirement. Their t_c^{γ} distribution is shown in Fig. 1. Our fit to the data outside the signal region predicts total backgrounds of 6.2 ± 3.5 from cosmic rays, 6.8 ± 4.9 from beam halo background sources, and the rest from the standard model with a measured wrong vertex fraction of (0.5 ± 0.2) %. Inside the signal region, {2, 10} ns, we predict 1.25 ± 0.66 events: 0.71 ± 0.60 from standard model, 0.46 ± 0.26 from cosmic rays, and 0.07 ± 0.05 from beam halo. Two events are observed in the data. Since the result is consistent with the no-signal hypothesis, we set predicted and observed limits on the $\hat{\chi}_1^0$ lifetime and mass. Figure 2 shows the contours of constant 95% C.L. cross section upper limit which reflects the change in acceptance. The 95% C.L. exclusion region, where the predicted production cross section at next-to-leading order [19] is



FIG. 3 (color online). The exclusion region at 95% C.L. as a function of χ_1^0 lifetime and mass for a GMSB model [9]. The predicted and the observed regions are shown separately and are compared to the most stringent published limit from LEP searches [8].

above the upper limit, is shown in Fig. 3, taking into account the uncertainties on the parton distribution functions (6%) and the renormalization scale (2%). These limits extend at large masses beyond those of LEP searches using photon "pointing" methods [8].

In conclusion, we have performed the first search for heavy, long-lived particles that decay to photons at a hadron collider using data collected with the EMTiming system at the CDF II detector. There is no excess of events beyond expectations. As our search strategy does not rely on event properties specific to GMSB models, we can exclude any delayed $\gamma + \text{jet} + \not{\!\!\!E}_T$ signal that would produce more than 5.5 events. We set cross section limits using a supersymmetric model with $\chi_1^0 \rightarrow \gamma \tilde{G}$, and find a GMSB exclusion region in the χ_1^0 lifetime vs mass plane with the world-best 95% C.L. lower limit on the χ_1^0 mass of 101 GeV/ c^2 at $\tau_{\chi_1^0} = 5$ ns. By the end of run II, an integrated luminosity of 10 fb⁻¹ is possible for which we estimate a mass reach of $\simeq 140 \text{ GeV}/c^2$ at a lifetime of 5 ns by scaling the expected number of background events.

We thank the Fermilab staff and the technical staffs of the participating institutions for their vital contributions. This work was supported by the U.S. Department of Energy and National Science Foundation; the Italian Istituto Nazionale di Fisica Nucleare; the Ministry of Education, Culture, Sports, Science and Technology of Japan; the Natural Sciences and Engineering Research Council of Canada; the National Science Council of the Republic of China; the Swiss National Science Foundation; the A.P. Sloan Foundation; the Bundesministerium für Bildung und Forschung, Germany; the Korean Science and Engineering Foundation and the Korean Foundation; the Particle Physics Research and Astronomy Research Council and the Royal Society, UK; the Russian Foundation for Basic Research: the Comisión

Interministerial de Ciencia y Tecnología, Spain; in part by the European Community's Human Potential Programme under Contract No. HPRN-CT-2002-00292; and the Academy of Finland.

- ^aVisiting scientist from the University of Athens.
- ^bVisiting scientist from the University of Bristol.
- ^cVisiting scientist from the University Libre de Bruxelles.
- ^dVisiting scientist from Cornell University.
- ^eVisiting scientist from the University of Cyprus.
- ^fVisiting scientist from the University of Dublin.
- ^gVisiting scientist from the University of Edinburgh.
- ^hVisiting scientist from the University of Heidelberg.
- ⁱVisiting scientist from the Universidad Iberoamericana.
- ^jVisiting scientist from the University of Manchester.
- ^kVisiting scientist from the Nagasaki Institute of Applied Science.
- ¹Visiting scientist from the University de Oviedo.
- ^mVisiting scientist from the University of London, Queen Mary College.
- ⁿVisiting scientist from the University of California, Santa Cruz, USA.
- ^oVisiting scientist from Texas Tech University, USA.
- ^pVisiting scientist from the University of California, Irvine, USA.
- ^qVisiting scientist from IFIC(CSIC-Universitat de Valencia).
- J. L. Feng, A. Rajaraman, and F. Takayama, Phys. Rev. D 68, 063504 (2003); M. J. Strassler and K. M. Zurek, arXiv:hep-ph/0605193.
- S. Ambrosanio *et al.*, Phys. Rev. D 54, 5395 (1996); C. H.
 Chen and J. F. Gunion, Phys. Rev. D 58, 075005 (1998).
- [3] F. Abe *et al.* (CDF Collaboration), Phys. Rev. Lett. **81**, 1791 (1998); Phys. Rev. D **59**, 092002 (1999).
- [4] D. Acosta *et al.* (CDF Collaboration), Phys. Rev. D 71, 031104 (2005).
- [5] V. Abazov *et al.* (D0 Collaboration), Phys. Rev. Lett. **94**, 041801 (2005).

- [6] P. Wagner, Ph.D. thesis, Texas A&M University, 2007.
- [7] D. Acosta *et al.* (CDF Collaboration), Phys. Rev. D 71, 032001 (2005).
- [8] A. Heister *et al.* (ALEPH Collaboration), Eur. Phys. J. C
 25, 339 (2002); also see M. Gataullin, S. Rosier, L. Xia, and H. Yang, arXiv:hep-ex/0611010; G. Abbiendi *et al.* (OPAL Collaboration), Proc. Sci., HEP2005 (2006) 346; J. Abdallah *et al.* (DELPHI Collaboration), Eur. Phys. J. C
 38, 395 (2005).
- [9] B.C. Allanach *et al.*, Eur. Phys. J. C 25, 113 (2002). We use benchmark model 8 and take the messenger mass scale M_{mess} = 2Λ, tan(β) = 15, μ > 0, and the number of messenger fields N_{mess} = 1. The G mass factor and the supersymmetry breaking scale Λ are allowed to vary independently.
- [10] D. Toback and P. Wagner, Phys. Rev. D 70, 114032 (2004).
- [11] The distribution of the $p\bar{p}$ collisions has a standard deviation of 30 cm and 1.3 ns in z_i and t_i , respectively.
- [12] M. Goncharov *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A 565, 543 (2006).
- [13] The standard requirement, $\chi^2_{CES} < 20$ [F. Abe *et al.* (CDF Collaboration), Phys. Rev. D **52**, 4784 (1995)], has been removed because there is evidence that it is inefficient for photons that arrive with large incident angles relative to the face of the detector.
- [14] A. Bhatti *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A 566, 375 (2006). We use corrected jets reconstructed with a cone of 0.7.
- [15] T. Sjöstrand *et al.*, Comput. Phys. Commun. **135**, 238 (2001). We use version 6.216.
- [16] We use the standard GEANT based detector simulation[R. Brun *et al.*, CERN Report No. CERN-DD/EE/84-1, 1987] and add a parametrized EMTiming simulation.
- [17] E. Boos, A. Vologdin, D. Toback, and J. Gaspard, Phys. Rev. D 66, 013011 (2002).
- [18] J. Conway, CERN Yellow Book Report No. CERN 2000-005, 2000, p. 247. We assume a flat prior in the production cross section.
- [19] W. Beenakker *et al.*, Phys. Rev. Lett. **83**, 3780 (1999).