# Accurate Electron Densities at Nuclei using Small Ramp-Gaussian Basis Sets

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#### Abstract

Electron densities at nuclei are difficult to calculate accurately with all-Gaussian basis sets because they lack an electron-nuclear cusp. The newly developed mixed ramp-Gaussian basis sets, such as R-31G, possess electron-nuclear cusps due to the presence of ramp functions in the basis. The R-31G basis set is a general purpose mixed ramp-Gaussian basis set modelled on the 6-31G basis set. The prediction of electron densities at nuclei using R-31G basis sets for Li–F outperforms Dunning, Pople and Jensen general purpose all-Gaussian basis sets of triple-zeta quality or lower and the cc-pVQZ basis set. It is of similar quality to the specialised pcJ-0 basis set which was developed with partial decontraction of core functions and extra high exponent *s*-Gaussians to predict electron density at the nucleus. These results show significant advantages in the properties of mixed ramp-Gaussian basis sets compared to all-Gaussian basis sets.

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# **1** Introduction

Standard ab initio quantum chemistry methods use all-gaussian basis sets, which have no electronnuclear cusp and thus struggle to accurately predict core-dependent properties. Electron densities at nuclei,  $\rho_0$ , is an important example of a core-dependent chemical property that is directly response for the Fermi contact term of the indirect spin-spin coupling constant (SSCC) and is also important in the fine structure of the hydrogen spectrum and for relativistic calculations (in the Darwin term). Some studies use the synonym "contact density" for this property.

Unlike traditional all-gaussian basis sets, the newly developed<sup>1</sup> mixed ramp-Gaussian basis sets do have a cusp due to the presence of a ramp function, which in the simplest case has the unnormalised form  $(1-r)^n$  for  $r \le 1$  and 0 otherwise. These new basis sets are thus expected to give superior results to similar sized all-Gaussian basis sets for  $\rho_0$ ; this paper investigates this hypothesis. Electron density at the nuclei is both an important property in itself and a straightforward, easily explored example of a core-related parameter.

The new R-31G basis set was produced by replacing the 6-fold contracted core basis function in 6-31G with a two-fold contracted basis function with one ramp and one gaussian. The R-31G basis set was shown to produce very similar chemistry (e.g. ionisation energies, atomisation energies etc.) to the 6-31G basis set for unrestricted Hartree-Fock calculations of molecules in the first-row-only G2 dataset. It was recently demonstrated<sup>2</sup> that integrals that arise using ramp basis functions can be evaluated efficiently and that UHF/R-31+G calculation timings are competitive with UHF/6-31+G calculation timings for large linear molecules. For the small systems investigated here, calculation times for UHF/R-31G and UHF/6-31G are similar.

Atomic units are used throughout this paper.

# 2 Basis Sets

# 2.1 All-Gaussian basis sets

General-purpose basis sets are designed to give accurate description of chemical energies; for this purpose, inflexible descriptions of the core are generally adequate. Common general-purpose basis sets discussed in this paper include the Pople (e.g. 6-31G<sup>3</sup> and 6-311G<sup>4</sup>), Dunning<sup>5</sup> (cc-pVnZ) and Jensen<sup>6,7</sup> (pc-n) basis sets.

However, accurate quantification of  $\rho_0$ , and hence the Fermi contact term in the indirect spinspin coupling constant, requires a more flexible and accurate description of the core.<sup>8–13</sup> The pcJ-n and ccJ-pVnZ basis sets were designed by Jensen<sup>12</sup> to reproduce indirect spin-spin coupling constants accurately. In particular, they contain additional very high exponent core Gaussian primitives and decontraction of the core basis functions. Our tests show that the pcJ-n basis sets are slower than their parent pc-n bases (by about a factor of 2-4), but significantly quicker than the higher quality pcJ-(n+1) basis set.

Our calculations show that pcS-n, ccS-pVnZ or cc-pCVnZ basis sets do *not* give significantly improved results for  $\rho_0$  compared to their pc-n/ cc-pVnZ parent basis sets. This is because the additional primitives are not very high exponent *s* primitives.

## 2.2 Mixed ramp-Gaussian basis sets

The R-31G and R-31+G basis sets are general-purpose basis sets that derive many of their properties from their 6-31G and 6-31+G parent basis sets and, in particular, our present implementation of R-31G and 6-31G have similar calculation times.<sup>2</sup>

The R-31G basis set was not optimised specifically for electron densities at the nuclei. In fact, the 'R' basis function was found by maximising its overlap with the original 6-fold contracted core basis function in 6-31G. Therefore, it is only the inherent properties of the ramps that will be responsible for superior values of  $\rho_0$  from R-31G, not a consequence of the method of optimisation.

We also consider for the first time the R1-31G basis set which is obtained from the R-31G

basis by decontracting the 1s core function. This gives additional core flexibility whilst retaining the same number of primitives.

# 2.3 Slater basis sets

Calculations involving Slater basis sets are difficult, but can be done for very small systems. There are results from previous investigations of atoms<sup>14</sup> for  $\rho_0$  with three Slater basis sets, VB1, VB2 and VB3.<sup>15</sup>

The composition of these Slater basis sets<sup>15</sup> for Li and Be is VB1 [5s,1p], VB2 [6s,2p,1d] and VB3 [7s,3p,2d,1f], while for B-Ne, the composition of these basis sets<sup>15</sup> is VB1 [5s,3p,1d], VB2 [6s,4p,2d,1f] and VB3 [7s,5p,3d,2f,1g]. All basis functions are uncontracted.

The R-31G and R1-31G basis sets are therefore of similar size to the VB1 basis, but do not include polarisation functions.

# **3** Results and Discussion

## 3.1 Method

We investigate  $\rho_0$  for a set of molecules at the Hartree-Fock level of theory and compare results from different basis sets to benchmark values: the exact result is known for H, for other atomic systems except B we use numerical HF calculations, <sup>16,17</sup> for B HF/VB3 results are used, while for all molecular systems results from HF/pcJ-4 were taken as the benchmark. We report the error  $\Delta \rho_0$ for each calculation compared to the benchmark value.

Calculations for all-gaussian basis sets were performed using Q-CHEM.<sup>18</sup> Basis sets not built into the standard Q-CHEM installation were sourced from the online basis set exchange library.<sup>19,20</sup>

All calculations involving mixed ramp-Gaussian basis sets were done using the recently developed program RAMPITUP,<sup>2</sup> a Fortran90 program which calculates integrals in a mixed ramp-Gaussian basis set and performs the HF SCF calculation. It is important to note that comparison to experiment is currently not the fairest way to assess these new basis sets because all calculations are at the Hartree-Fock (HF) level. Thus, we have chosen to compare our results to highly accurate HF calculations; either numerical HF results, where available, or HF/pcJ-4 otherwise.

3.2 Atoms

#### 3.2.1 R-31G, R-31+G and R1-31G vs. Slater basis sets.

Since Slater and ramp basis functions both have non-zero electron-nuclear cusps, we quantify not only the electron density at the nucleus,  $\rho_0$ , but also quantify the ratio, G, between  $\rho_0$  and its derivative  $\rho'_0$  given by

$$G = -\frac{\rho_0'}{Z\rho_0},\tag{1}$$

where *Z* is the nuclear charge. It can be shown that for the exact wavefunction, G = 2, the so-called Kato-cusp condition.<sup>21</sup> Note that for an all-Gaussian basis set, G=0.

Tables 1 and 2 show that R-31G performs significantly worse (generally more than a factor of 4) than the three Slater basis sets in predicting the cusp properties  $\rho_0$  and *G*, despite both ramps and Slater functions having non-zero nuclear-electron cusps. R-31+G performs marginally better than R-31G, but this is a small effect and we will thus not consider the R-31+G basis set further in this paper. The inferior performance of the R-31G and R-31+G basis sets compared to the all-Slater basis set is somewhat disappointing, though the mixed ramp-Gaussian basis sets will of course still vastly outperform all-Gaussian basis sets in predicting *G*.

For  $\rho_0$ , the errors of R1-31G are about one-third the size of the R-31G errors and are comparable in quality to VB1 for heavy elements, though significantly worse for light elements. R1-31G overestimates  $\rho_0$ , whereas almost all other results underestimate this quantity. This occurs because the variational optimisation of the wavefunction preferences occupation of the "R" basis function to lower the overall energy of the system.

	Н	Li	Be	В	С	N	0	F	
Benchmark $\overline{\rho_0}$									
	0.318	13.83	35.43	71.98	127.56	206.13	311.97	448.71	
Large unpolarised basis set $\Delta  ho_0$									
unpol-pcJ-4	-0.000			-0.09	-0.12	-0.25	-0.36	-0.43	
Slater basis s	sets $\Delta \rho_0$								
VB1	0.000	0.00	-0.02	-0.07	-0.30	-0.63	-1.24	-1.54	
VB2	0.000	0.00	-0.02	-0.01	-0.14	-0.19	-0.45	-0.69	
VB3	0.000	0.00	-0.01	0.00	-0.09	-0.14	-0.32	-0.49	
Mixed ramp-Gaussian basis sets $\Delta  ho_0$									
R-31G	-0.020	-0.30	-0.72	-1.39	-2.35	-3.62	-4.91	-6.04	
R-31+G	-0.020	-0.28	-0.77	-1.44	-2.38	-3.63	-4.89	-5.99	
R1-31G	-0.020	0.35	0.43	0.55	0.74	0.95	1.32	1.93	
General-pur	pose all-Ga	aussian b	asis sets $\Delta$	$ ho_0$					
6-31G	-0.020	-0.90	-2.02	-4.22	-7.86	-13.04	-19.68	-27.71	
6-311G	-0.031	-0.79	-1.99	-3.86	-6.52	-10.48	-15.52	-21.73	
pc-0	-0.077	-2.20	-5.66	-11.47	-20.34	-32.87	-49.74	-71.41	
cc-pVDZ	-0.050	-0.63	-1.52	-3.07	-5.42	-8.76	-13.36	-19.19	
pc-1	-0.035	-1.15	-3.05	-6.26	-11.21	-18.20	-27.72	-39.86	
cc-pVTZ	-0.031	-0.32	-1.01	-2.82	-4.90	-7.85	-11.74	-16.75	
pc-2	-0.016	-0.39	-1.23	-2.53	-4.42	-7.28	-10.87	-15.55	
cc-pVQZ	-0.020	-0.31	-0.70	-1.40	-2.47	-4.09	-6.07	-8.80	
pc-3	-0.007	-0.01	-0.18	-0.46	-0.98	-1.76	-2.79	-4.06	
Specialised all-Gaussian basis sets $\Delta ho_0$									
pcJ-0	-0.012			-1.30	-2.25	-3.62	-5.55	-7.90	
ccJ-pVDZ	-0.004			-0.33	-0.61	-1.01	-1.54	-2.13	
pcJ-1	-0.005			-0.79	-1.38	-2.20	-3.29	-4.57	
ccJ-pVTZ	-0.030			-0.40	-0.54	-0.89	-1.32	-1.79	
pcJ-2	-0.002			-0.36	-0.63	-1.01	-1.48	-1.97	
ccJ-pVQZ	-0.002			-0.17	-0.33	-0.56	-0.81	-1.07	
pcJ-3	0.000			0.03	0.03	0.01	0.06	0.19	

Table 1: The first row shows benchmark for  $\rho_0$  for atoms; subsequent rows show errors in  $\rho_0$  using different basis sets.

	Li	Be	В	С	Ν	0	F			
Mixed ramp-Gaussian basis sets										
R-31G	-0.086	-0.06	-0.05	-0.04	-0.04	-0.03	-0.03			
R-31+G	-0.085	-0.06	-0.05	-0.04	-0.04	-0.03	-0.03			
R1-31G	-0.061	-0.05	-0.04	-0.03	-0.03	-0.02	-0.02			
All-Slater basis sets										
VB1	0.010	0.01	0.00	-0.00	-0.01	-0.01	-0.01			
VB2	0.011	0.01	0.01	0.00	0.00	0.00	-0.00			
VB3	0.011	0.01	0.01	0.00	0.00	0.00	0.00			

Table 2: Errors in the HF cusp ratio,  $G = -\rho'(0)/(Z\rho_0)$ . Results for Slater basis sets from ref 14. The Kato cusp condition gives the exact benchmark result as G = 2.

#### 3.2.2 R-31G and R1-31G vs. common all-Gaussian basis sets.

By basis set definition, R-31G, R1-31G and 6-31G results for atomic hydrogen are identical. In Table 1 we observe that for hydrogen  $\rho_0$  results from these basis sets are surprisingly good (almost the same as cc-pVQZ).

For non-hydrogen atoms, Table 1 shows that in absolute terms, R-31G only underestimates  $\rho_0$  by 2%. This is a better prediction than cc-pVQZ, and about half the quality of the pc-3 calculation. This is a very promising result; R-31G is a much smaller basis set than the quadruple zeta all-Gaussian basis sets, yet produces comparable  $\rho_0$  values. This result is attributed to the non-zero cusp on the ramp in R-31G which is inherently able to model  $\rho_0$  more accurately than all-Gaussian basis sets. The R1-31G basis set performs even better, with errors in line with pc-3 results.

To put this result in context, the one-electron Darwin energy for an atom is given by  $E_{\text{Darwin}} = \pi \alpha^2 Z \rho_0 E_h = 0.167294 Z \rho_0 mE_h$ , where  $\alpha$  is the fine structure constant. For carbon, this equates to 128 mE<sub>h</sub>(336 kJ/mol); 6-31G underestimates this energy by 7.9 mE<sub>h</sub>(21 kJ/mol), R-31G by 2.4 mE<sub>h</sub>(6.2 kJ/mol) and R1-31G by 0.74 mE<sub>h</sub>(2.0 kJ/mol).

#### 3.2.3 R-31G and R1-31G vs. specialised Gaussian basis sets.

In Table 1, for elements B-F, we observe that R-31G prediction of  $\rho_0$  are of approximately the same quality as those for the larger specialised basis set pcJ-0 (slightly worse for B- N and slightly better for O-F). The R-31G results are worse than for this larger specialised basis set.

Increasing the flexibility of the basis set yields clear dividends; with only one extra basis function and no extra primitives, R1-31G predictions are clearly better than pcJ-0 and pcJ-1 results, and comparable to ccJ-pVDZ, ccJ-pVTZ and pcJ-2 results. R1-31G is a much smaller basis set and calculations already are much faster than with larger all-Gaussian basis sets.

## 3.3 Molecules

Electron density at the nucleus is a core property and affected relatively little by the molecular environment of that nucleus. Therefore, the overall quality of predictions of  $\rho_0$  by different basis sets in molecules will be similar to the quality of results in atoms. This is quantified by the mean error in the electron density at the nuclei across the test set,  $\Delta \rho_0$ .

In molecules, it is also interesting to see how well a basis set can describe the changes in electron density at the nucleus as the molecular environment varies. This is quantified by the standard deviation (SD) of  $\Delta \rho_0$ . If the SD of  $\Delta \rho_0$  is small, then  $\rho_0$  for the molecule in the basis set can be corrected by a single number (i.e. the error in  $\rho_0$  is systematic). However, if the SD of  $\Delta \rho_0$  is large, then a systematic correction factor cannot be used; this indicates that the description of the electron density near the nuclei is not sufficiently flexible to accurately describe the effect of varying molecular environment on  $\rho_0$ .

#### 3.3.1 Hydrogen nuclei.

Though it is possible in molecules for 6-31G, R-31G and R1-31G to give different results for  $\rho_0$  through an indirect effect of the modified basis set of nearby heavy atoms, the full set of molecular results show this difference is negligible in practice. As summarised in Table 3, 6-31G, R-31G

Table 3: The first row shows the mean and standard deviation of  $\rho_0$  values for our test sets. Subsequent rows show the mean and standard deviation of  $\Delta \rho_0$ . The test sets contains the following molecules: CH,  ${}^{3}CH_2$ , CH<sub>3</sub>, CH<sub>4</sub>, C<sub>2</sub>H<sub>2</sub>, C<sub>2</sub>H<sub>4</sub>, C<sub>2</sub>H<sub>6</sub>, CN, HCN, CH<sub>3</sub>OH, CO, HCO, CO<sub>2</sub>, NH, NH<sub>2</sub>, NH<sub>3</sub>, N<sub>2</sub>, NO, O<sub>2</sub>, OH, H<sub>2</sub>O, H<sub>2</sub>O<sub>2</sub>. All geometries are MP2/6-31G(d).

	Н		С		N		0			
	Mean	SD	Mean	SD	Mean	SD	Mean	SD		
Benchmark $\rho$	0									
pcJ-4	0.463	0.022	126.21	0.50	205.09	0.50	310.56	0.39		
Large unpola	rised basis									
unpol-pcJ-4	-0.001	0.003	0.03	0.04	0.04	0.03	0.07	0.05		
Mixed ramp-Gaussian basis sets $\Delta  ho_0$										
R-31G	-0.045	0.011	-2.13	0.12	-3.14	0.12	-4.16	0.09		
R1-31G	-0.045	0.011	0.91	0.04	1.37	0.05	1.99	0.07		
General-purpose all-Gaussian basis sets $\Delta  ho_0$										
6-31G	-0.045	0.011	-7.61	0.11	-12.56	0.12	-18.90	0.10		
6-311G	-0.038	0.014	-6.30	0.04	-10.09	0.04	-14.91	0.04		
pc-0	-0.124	0.110	-20.30	0.14	-32.89	0.10	-49.75	0.12		
cc-pVDZ	-0.096	0.006	-5.15	0.10	-8.43	0.11	-12.83	0.08		
pc-1	-0.086	0.005	-10.98	0.08	-17.91	0.07	-27.21	0.07		
cc-pVTZ	-0.051	0.004	-4.76	0.03	-7.60	0.03	-11.41	0.02		
pc-2	-0.035	0.002	-4.31	0.06	-7.10	0.02	-10.56	0.05		
cc-pVQZ	-0.031	0.002	-2.33	0.01	-3.82	0.01	-5.66	0.01		
pc-3	-0.014	0.001	-0.94	0.03	-1.55	0.01	-2.44	0.01		
Specialised all-Gaussian basis sets $\Delta  ho_0$										
pcJ-0	-0.008	0.012	-2.44	0.12	-3.87	0.10	-5.85	0.14		
ccJ-pVDZ	-0.003	0.004	-0.47	0.01	-0.75	0.01	-1.14	0.01		
pcJ-1	-0.002	0.004	-1.24	0.01	-1.96	0.01	-2.91	0.01		
ccJ-pVTZ	-0.003	0.001	-0.40	0.00	-0.63	0.00	-0.94	0.00		
pcJ-2	-0.001	0.000	-0.50	0.00	-0.76	0.00	-1.10	0.00		
ccJ-pVQZ	-0.001	0.000	-0.20	0.00	-0.31	0.00	-0.44	0.00		
pcJ-3	0.001	0.000	0.15	0.00	0.26	0.00	0.42	0.00		

and R1-31G gives surprisingly good results for the absolute value of  $\rho_0$  (about equivalent to ccpVTZ). However, they do not represent the change in  $\rho_0$  for different molecular species as well as cc-pVTZ; this is quantified by the SD of  $\Delta \rho_0$  (0.011 for X-31G vs. 0.004 for cc-pVTZ).

#### 3.3.2 Non-hydrogen nuclei.

The first two rows in Table 3 illustrate the importance of polarisation functions for calculating  $\rho_0$ ; the large unpolarised basis set *unpol*-pcJ-4 is less able to describe the variation of  $\rho_0$  between different molecules than the fully polarised basis set (quantified by the SD of  $\Delta \rho_0$ ). Therefore, the absence of polarisation functions in R-31G will immediately reduce its ability to describe the variation of  $\rho_0$  between different molecules (i.e. increase the SD of  $\Delta \rho_0$ ) compared to fullypolarised basis sets.

In Table 3, we observe the R-31G predictions of the absolute value of  $\rho_0$  (quantified by the mean of  $\Delta \rho_0$ ) for C, N, O and F nuclei in molecules, like in atoms, outperforms predictions of all general-purpose triple-zeta basis sets and cc-pVQZ, while being worse than the pc-3 basis set by about a factor of 2. In absolute terms, R-31G generally underestimates  $\rho_0$  by about 1-2%.

Table 3 shows that R-31G outperforms the slightly larger pcJ-0 basis set by about 20%, but is not as good as all the higher-quality specialised spin-spin coupling basis sets. It is interesting to consider the carbon atom specifically. In atomic carbon, the pcJ-0 result is better than the R-31G result, whereas for molecular systems, the R-31G result is superior. This is a secondary effect of pcJ-0's inferior description of the valence region on the core electrons.

The errors from R-31G and pcJ-0 between different molecules are not as systematic as for larger basis set, such as cc-pVQZ, as quantified by the SD of  $\Delta \rho_0$ . Therefore, the cc-pVQZ errors are more predictable (and thus easier to correct and more likely to cancel when considering relative differences) than the R-31G or pcJ-0 basis set errors. This can be attributed to the small basis sets' lack of flexibility.

R1-31G increases the core basis set flexibility of R-31G and this shows significant benefit in reducing the SD of  $\Delta \rho_0$  by about two-thirds (e.g. from 0.12 to 0.04 for carbon nuclei). The SD of

 $\Delta \rho_0$  for R1-31G are close to those of the unpolarised basis set limit benchmark; significant further improvements probably require the introduction of polarisation functions. It is encouraging that a simple decontraction of the core function, resulting in only one additional basis function and no additional primitives (and thus no extra integrals), yields a basis with most of the flexibility of the unpolarized benchmark basis. By way of illustration, the primitive and contracted basis function composition of R1-31G is (6s4p)  $\rightarrow$  [4s2p], whereas for the unpolarised pcJ-4 basis it is (19s12p)  $\rightarrow$  [11s9p]. Furthermore, the SD of  $\Delta \rho_0$  for R1-31G is roughly in line with pc-2 and three times better than pcJ-0 (0.04 vs. 0.06 vs. 0.12 for carbon nuclei).

The SD of  $\Delta \rho_0$  will equate to an error in the properties between the atom and molecule. To provide context, we can convert this SD to an average error in the Darwin energy per carbon atom in molecules compared to the Darwin energy of the carbon atom. For nitrogen, 6-31G and R-31G will give errors of 0.12 mE<sub>h</sub> (0.32 kJ/mol), R1-31G gives errors of 0.05 mE<sub>h</sub> (0.13 kJ/mol) while cc-pVQZ errors are 0.01 mE<sub>h</sub> (0.03 kJ/mol). It is clear, then, why we can ignore this error in the Darwin energy in calculating atomisation energies for light elements, while it will be nonnegligible for very high precision calculations or for heavier atoms.

Based on these results, we recommend R1-31G as a cheap, high accuracy method of calculating the Fermi contact term in indirect spin-spin coupling constants and for calculating Darwin relativistic corrections. Addition of polarisation functions is expected to improve the performance of this basis set.

# 4 Conclusions

For non-hydrogen nuclei, despite its small size, the R-31G basis set produces very good results for HF electron densities at the nucleus because the ramp has a non-zero derivative at the origin. The small, general-purpose mixed ramp-Gaussian basis set R-31G outperforms Dunning, Jensen and Pople general-purpose all-Gaussian basis sets of up to triple-zeta quality as well as cc-pVQZ. R-31G has about equivalent performance to pcJ-0 (a small specialised basis set). However, it is worse than pc-3 and all specialized basis sets of double-zeta quality and higher. Decontracting the ramp and Gaussian primitive in the core basis function to produce the R1-31G gives even better performance, reducing absolute errors ( $\Delta \rho_0$ ) by about two-thirds.

R-31G has very little core flexibility and is thus significantly worse than cc-pVQZ at reproducing the changes in  $\rho_0$  in different molecules (evidenced by the larger standard deviation of the  $\rho_0$ errors in molecules). The more flexible R1-31G enhances the ability to describe variations in  $\rho_0$ substantially, leading to lower SD of  $\Delta \rho_0$ ; about equivalent to pc-2. To significantly improve the performance of mixed ramp-Gaussian basis sets, polarisation basis functions need to be introduced to the basis set; this is an important future area of development and a key motivator for extending the capacities of the RAMPITUP program to higher angular momentum Gaussians.

These results show an application (electron density at the nucleus) where the superior cusp and inner-electron behaviour of the mixed ramp-Gaussian basis sets gives significant advantage over all-Gaussian basis sets. These results provides additional justification for the development of a fully-integrated ramp-Gaussian integral package in common quantum chemistry programs. The development of specialised and larger mixed ramp-Gaussian basis sets in the future will be extremely beneficial in further exploring the ability of ramp basis functions to model the electron density at the nuclei and other core-dependent properties.

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