## **1** Diamonds from the lower mantle?

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5 Natural diamonds, because of their great physical resiliency, can preserve information about their 6 formation, storage and transport conditions for billions of years. Diamond samples therefore 7 provide a unique opportunity to directly study ancient samples of the Earth's deep interior. In order to correctly interpret the information diamonds provide, it is essential to accurately constrain the 8 9 depth of their origin. This depth provenance is usually identified using coexisting minerals, which are occasionally trapped as inclusions within diamonds during their growth. Comparison of an 10 11 inclusion's composition and mineralogy with experimental phase equilibria allows the diamond's growth conditions to be estimated. While the majority of diamonds likely originate from depths of 12 13 140-220 km in cratonic mantle, a small subset appears to have been exhumed from depths extending to > 800 km, called "superdeep" or "ultradeep" diamonds (Walter et al. 2011; Pearson et 14 15 al. 2014). Inclusions of magnesiowüstite are among the most commonly described in sub-16 lithospheric diamonds, and have often been assumed to indicate diamond provenance in the lower mantle because [Mg,Fe]O is not stable at upper mantle conditions in a subsolidus mantle 17 compositions (Trønnes 2009). This is despite the stability field of [Mg,Fe]O extending to ambient 18 19 pressure conditions and experimental evidence of magnesiowüstite stability in equilibrium with 20 diamond throughout the upper mantle (Brey et al. 2004; Thomson et al. 2016). A new study by 21 Uenver-Thiele et al. (2017) in American Mineralogist places important new constraints on the 22 formation and uplift history of inclusions containing magnesioferrite.

23 Studies of magnesiowüstite inclusions in diamonds from the Juina region of Brazil often report 24 observation of nanometre-sized crystals of magnesioferrite ( $[Mg, Fe^{2+}]Fe^{3+}_{2}O_{4}$ ), which supposedly 25 "confirm" the lower mantle origin of these samples. The magnesioferrite precipitates can occur at the interface between the diamond and [Mg,Fe]O inclusion, or as evenly distributed dislocation 26 27 "necklaces" within the inclusion interior (Harte et al. 1999; Wirth et al. 2014; Palot et al. 2016). Wirth et al. (2014) describe chains of globular [Mg<sub>0.5</sub>Fe<sub>0.5</sub>]Fe<sub>2</sub>O<sub>4</sub> crystals, ~ 75 nm in size, making 28 29 up 6-11 vol.% of the entire [Mg<sub>27</sub>Fe<sub>71</sub>]O inclusion. This suggests the original inclusion had an Fe<sup>3+</sup>/ $\Sigma$ Fe of 11-14 %, compared with 7 ± 2 % in the recovered magnesiowüstite (McCammon 30 1997). Wirth et al. (2014) also identified the magnesioferrite is accompanied by small, ~ 10-30 nm, 31 32 cubic voids, Al-bearing spinel and Ni-Fe metal blebs. Palot et al. (2016) describe isolated 10-20 nm octahedra of Mg[Fe<sub>0.75</sub>Cr<sub>0.17</sub>Al<sub>0.08</sub>]<sub>2</sub>O<sub>4</sub> throughout a [Mg<sub>84</sub>Fe<sub>16</sub>]O host with a recovered Fe<sup>3+</sup>/ $\Sigma$ Fe 33 34 content of 1-2 % that also contains ~ 30 ppm H<sub>2</sub>O in brucite precipitates. The bulk inclusion

35 composition reported by Palot et al. (~ [Mg72Fe28]O ignoring minor elements) implies the original magnesiowüstite must have had an Fe<sup>3+</sup>/ $\Sigma$ Fe of approximately 10-12 %. Wirth et al. (2014) and 36 Palot et al. (2016) both observe a topotaxial relationship between magnesioferrite lamellae and the 37 38 [Mg,Fe]O host, confirming the magnesioferrite must have formed during exsolution from a 39 homogenous magnesiowüstite grain. Using different arguments both studies concluded that the 40 magnesioferrite lamellae are indicative of the lower mantle provenance of these diamonds. Wirth et 41 al. (2014) suggested the highly non-stoichiometric magnesiowüstite inclusion sampled the highspin-low-spin transition in the in  $\varepsilon$ -iron stability field, promoting high Fe<sup>3+</sup> contents. This would 42 place inclusion, and diamond, formation near the very base of the mantle. Alternatively Palot et al. 43 44 (2016) interpretted the conditions of magnesioferrite exsolution using a phase diagram constructed 45 from atmospheric-pressure experimental data in the MgO-Fe<sub>2</sub>O<sub>3</sub>, MgO-Al<sub>2</sub>O<sub>3</sub> and MgO-Cr<sub>2</sub>O<sub>3</sub> 46 systems. This approach suggested that the onset of exsolution occurred at a temperature of ~ 1700 47 °C, which corresponds to ~ 25 GPa on the mantle adiabat (Palot et al. 2016). Both approaches 48 makes many assumptions and lack experimental verification that magnesioferrite exsolution 49 unambiguously indicates a diamond exhumation from the lower mantle. Indeed, as outlined below, the high ferric iron contents of the inclusions and new phase relations of magnesioferrite (Uenver-50 51 Thiele et al. 2017) instead point to a much shallower origin.

At low pressures (< 5 GPa) it is well understood that magnesiowüstite can incorporate significant 52 ferric iron, up to  $Fe^{3+}/\Sigma Fe$  of 70 %, mainly charge balanced by negative cation vacancies (e.g. 53 Hazen and Jeanloz 1984; Dobson et al. 1998). With increasing pressure and decreasing oxygen 54 55 fugacity the ferric iron capacity of magnesiowüstite decreases, due to a high-pressure phase transition of Fe<sub>3</sub>O<sub>4</sub> (Huang and Bassett 1986; McCammon et al. 1998). Since the mantle becomes 56 57 more reduced with depth, from ~ 1 log unit above the nickel-nickel oxide buffer (NNO+1) at 200 58 km to 1.5 log units below the iron-wüstite buffer (IW-1.5) at 660 km (Rohrbach and Schmidt 2011), 59 it is expected that ferric iron concentration of [Mg,Fe]O will fall rapidly with increasing formation 60 pressure. Indeed experiments confirm at conditions just within the lower mantle the maximum  $Fe^{3+}/\Sigma Fe$  in [Mg<sub>70</sub>Fe<sub>30</sub>]O, similar in composition to the inclusion observed by Palot et al. (2016), is 61 62 < 2% at NNO and < 0.5% at IW (Otsuka et al. 2013). Similarly [Mg<sub>20</sub>Fe<sub>80</sub>]O, similar to that observed by Wirth et al. (2014), would have a Fe<sup>3+</sup>/ $\Sigma$ Fe capacity of ~ 7 - 14 % at IW and NNO 63 64 respectively. These ferric iron capacities provide an upper bound, because "normal" lower mantle 65 conditions are more reduced and extend to higher pressure than the experimental conditions. Thus, 66 the bulk composition of diamond-hosted inclusions displaying magnesioferrite exsolution appears 67 inconsistent with formation under lower mantle conditions, unless exceptionally oxidsed conditions 68 are present.

69 In this issue, Uenver-Thiele et al. (2017) experimentally determined the high-pressure phase 70 relations of magnesioferrite (MgFe<sub>2</sub>O<sub>4</sub>) using the multi anvil apparatus. Prior to this study it was believed that MgFe<sub>2</sub>O<sub>4</sub> had a relatively simple phase diagram, with the ambient cubic spinel 71 72 structure (Fd-3m) stable until an isochemical phase transition to orthorhombic CaMn<sub>2</sub>O<sub>4</sub> structure 73 (Pbcm), HP- MgFe<sub>2</sub>O<sub>4</sub>, at ~ 17 GPa and temperatures above 1700 °C, or breakdown to MgO + 74 Fe<sub>2</sub>O<sub>3</sub> at lower temperatures (Levy et al. 2004). This chemography makes the interpretations of 75 Wirth et al. (2014) and Palot et al. (2016) feasible. However, the experiments of Uenver-Thiele et 76 al. (2017) have revealed a very different phase diagram, where the spinel-structured MgFe<sub>2</sub>O<sub>4</sub> 77 decomposes at ~ 10 GPa. It forms a phase assemblage of MgO +  $Fe_2O_3$  at temperatures below 1200 78  $^{\circ}$ C or Fe<sub>2</sub>O<sub>3</sub> + an unrecoverable phase of Mg<sub>5</sub>Fe<sub>2</sub>O<sub>8</sub>-Mg<sub>4</sub>Fe<sub>2</sub>O<sub>7</sub> stoichiometry at higher 79 temperatures. At pressures beyond  $\sim 13$  GPa the unrecoverable phase(s) are replaced by 80 orthorhombic, CaFe<sub>3</sub>O<sub>5</sub> structured (*Cmcm*), Mg<sub>2</sub>Fe<sub>2</sub>O<sub>5</sub> (Boffa Ballaran et al. 2015). HP-MgFe<sub>2</sub>O<sub>4</sub> 81 was not observed at any conditions up to 18 GPa and 1300 °C in this study. Further high-pressure 82 experiments are required in order to determine the structure(s) of the unrecoverable phase(s) using 83 in-situ methods, the full extent of the Mg<sub>2</sub>Fe<sub>2</sub>O<sub>5</sub> stability field and whether HP-MgFe<sub>2</sub>O<sub>4</sub> becomes 84 stable at higher pressures as suggested by previous studies (Andrault and Bolfan-Casanova 2001; 85 Levy et al. 2004).

The phase relations determined by Uenver-Thiele et al. (2017), coupled with the low ferric iron 86 87 capacity of magnesiowüstite in the lower mantle, have very significant consequences for the interpretation diamond formation pressures. Firstly, magnesioferrite is not stable at lower mantle 88 89 conditions where the diamond inclusions (Wirth et al. 2014; Palot et al. 2016) were believed to have 90 formed. Secondly, if the magnesioferrite did exsolve from (Mg,Fe)O as HP-MgFe<sub>2</sub>O<sub>4</sub> in the lower 91 mantle, it could not have directly inverted to the spinel structure, due to the large stability field of 92  $Mg_2Fe_2O_5 + Fe_2O_3$  as previously suggested. The presence of an additional minor phase between the 93 magnesioferrite platelets (Wirth et al. 2014) does suggest the magnesioferrite results from inversion 94 of lamellae of alternative stoichiometry. This idea that magnesioferrite resulted from the conversion 95 of Mg<sub>2</sub>Fe<sub>2</sub>O<sub>5</sub> + Fe<sub>2</sub>O<sub>3</sub> into magnesioferrite at  $\sim$  300 km depth requires further investigation. 96 However, the phase relations determined by Uenver-Thiele et al. (2017) demonstrate that 97 magnesioferrite exsolution from magnesiowüstite is not an indicator of formation in the lower mantle. Instead, it suggests a maximum depth for exsolution of ~ 10 GPa. While the conditions of 98 99 original inclusion entrapment of the samples described previously (Wirth et al. 2014; Palot et al. 100 2016) remain uncertain without further studies, the high ferric iron contents and magnesioferrite 101 phase relations are consistent with formation in the upper mantle or transition zone, possibly from oxidized slab materials. The study of Uenver-Thiele et al. (2017) highlights the potentially rich and 102

- 103 unexplored chemography and importance of post-spinel phase relations for understanding the
- 104 Earth's fundamental geochemical and geodynamic cycles.
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